

CMS Experiment at LHC, CERN
 Data recorded: Mon May 28 01:16:20 2012 CEST
 Run/Event: 195099 / 35438125
 Lumi section: 65
 Orbit/Clock: 16960411 / 2895

WHAT MAKES US THINK THAT PHYSICS BEYOND THE STANDARD MODEL EXISTS?



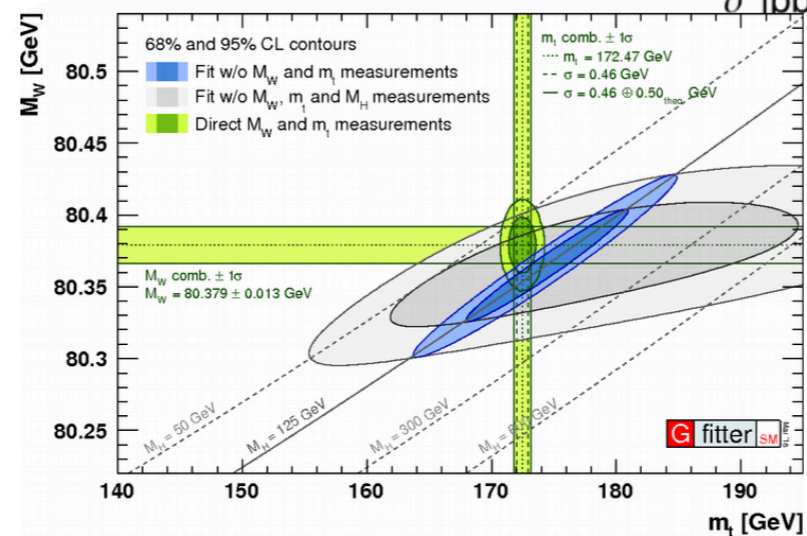
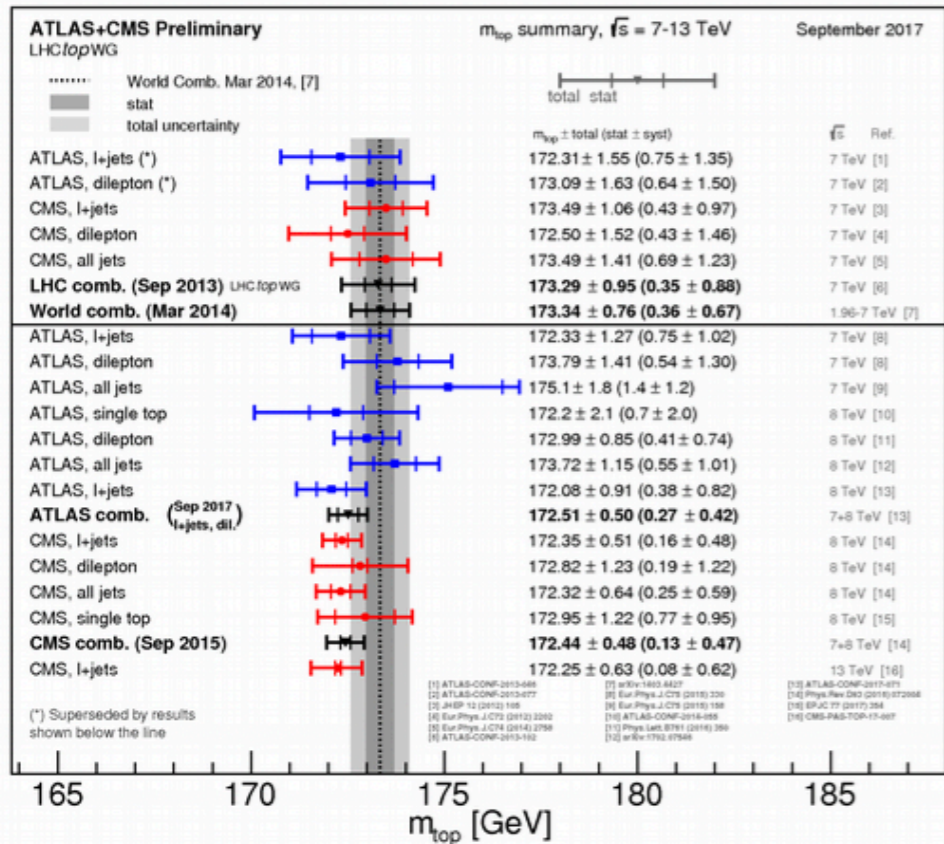
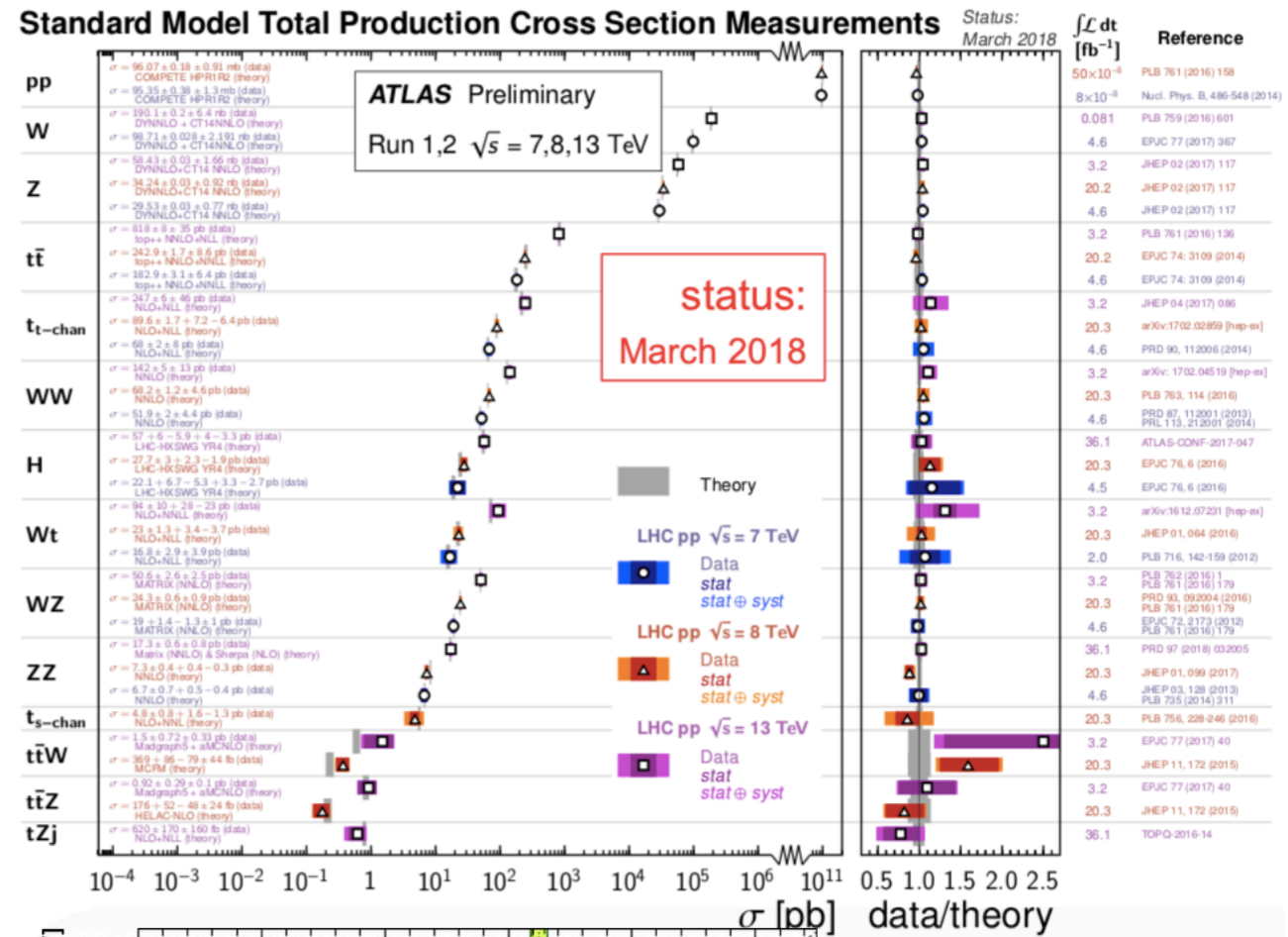
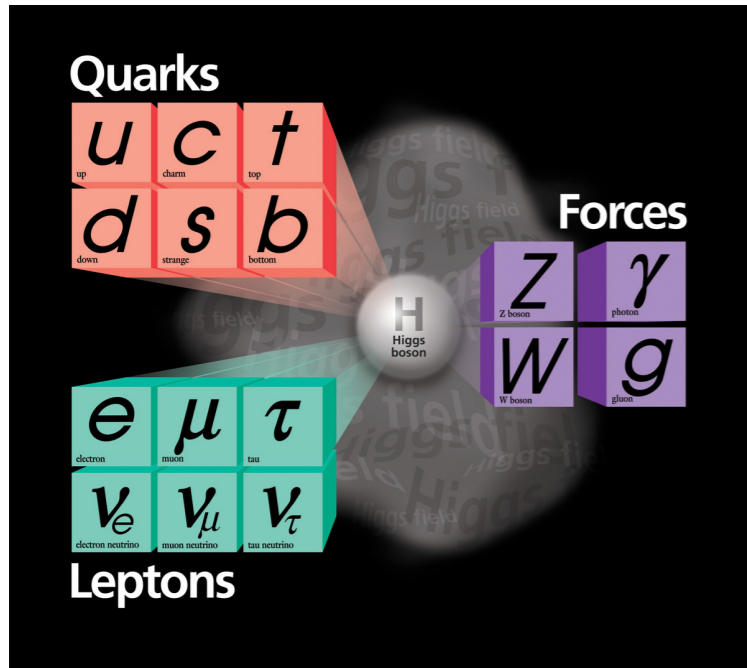
$E_T \sim 2 \text{ TeV}$
 with $E_T > 40 \text{ GeV}$
 Estimated $PU \sim 50$

Dmitry Kazakov
 BLTP JINR



THE STANDARD MODEL: THE STATUS REPORT AND OPEN QUESTIONS

The Standard Model



Extraordinary agreement between measurements and SM predictions

The Standard Model of Fundamental Interactions

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graph TD; A[The Standard Model of Fundamental Interactions] --> B[Higgs Sector]; A --> C[Neutrino Sector]; A --> D[Flavour Sector]; B --> E[New particles and Interactions]; C --> E; D --> E; C --> F[Dark Matter];
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Higgs Sector

Neutrino Sector

Flavour Sector

Dark Matter

New particles and Interactions

THE PRINCIPLES

- Three gauged symmetries $SU(3) \times SU(2) \times U(1)$
- Three families of quarks and leptons (3 \times 2, 3 \times 1, 1 \times 2, 1 \times 1)
- Brout-Englert-Higgs mechanism of spontaneous EW symmetry breaking \rightarrow Higgs boson
- CKM and PMNS mixing of flavours
- CP violation via phase factors
- Confinement of quarks and gluons inside hadrons
- Baryon and lepton number conservation
- CPT invariance \rightarrow existence of antimatter

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The ST principles allow:

- Extra families of quarks and leptons
- Presence or absence of right-handed neutrino
- Majorana or Dirac nature of neutrino
- Extra Higgs bosons

THE OPEN QUESTIONS

Why's?

- 📌 why the $SU(3) \times SU(2) \times U(1)$?
- 📌 why 3 generations ?
- 📌 why quark-lepton symmetry?
- 📌 why V-A weak interaction?
- 📌 why L-R asymmetry?
- 📌 why B & L conservation?
- 📌 etc

How's?

- 📌 how confinement actually works ?
- 📌 how the quark-hadron phase transition happens?
- 📌 how neutrinos get a mass?
- 📌 how CP violation occurs in the Universe?
- 📌 how to protect the SM from would be heavy scale physics?

THE OPEN QUESTIONS

Why's?

- why the $SU(3) \times SU(2) \times U(1)$?
- why 3 generations ?
- why quark-lepton symmetry?
- why V-A weak interaction?
- why L-R asymmetry?
- why B & L conservation?
- etc

- Is it self consistent ?
- Does it describe all experimental data?
- Are there any indications for physics beyond the SM?
- Is there another scale except for EW and Planck?
- Is it compatible with Cosmology? Where is dark matter?

How's?

- how confinement actually works ?
- how the quark-hadron phase transition happens?
- how neutrinos get a mass?
- how CP violation occurs in the Universe?
- how to protect the SM from would be heavy scale physics?



- With the Higgs Boson discovery the Standard Model is completed !
- Why are we not satisfied and think that new physics exists and new discoveries will come?



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- Why are we not satisfied and think that new physics exists and new discoveries will come?

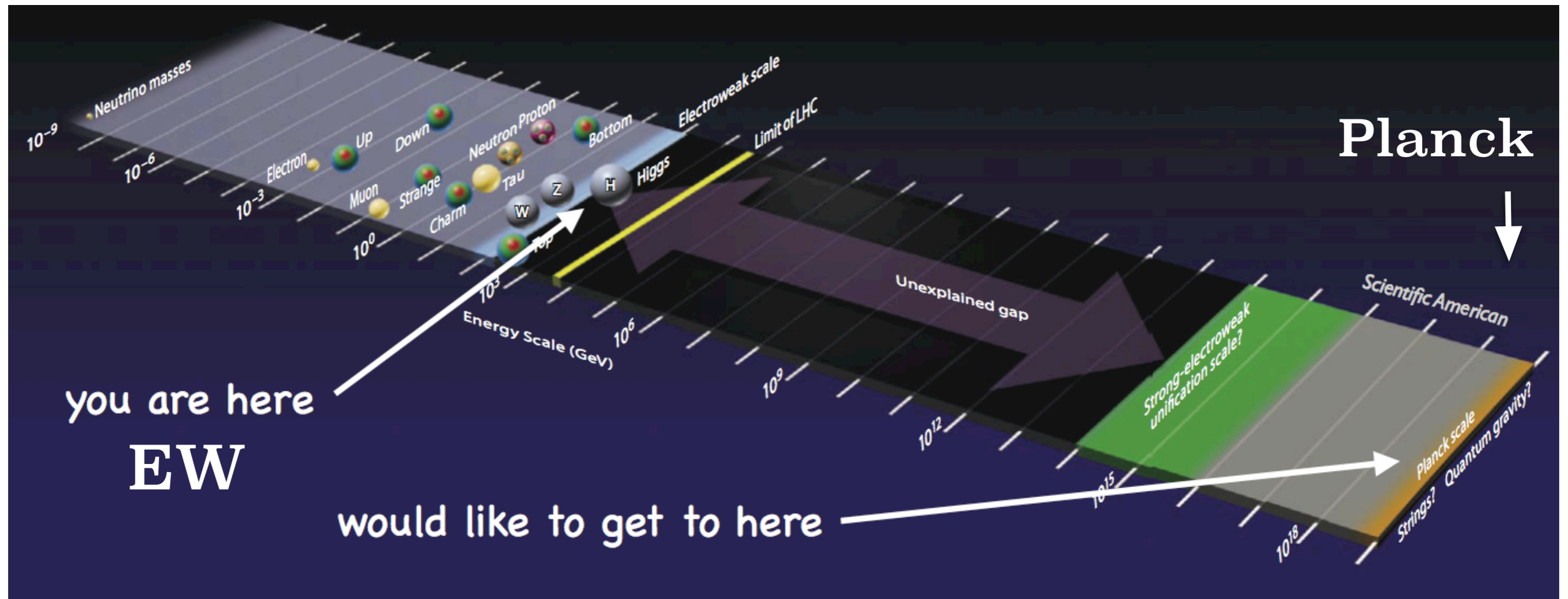


- There are conceptual problems which require a critical view beyond the SM
- There are small discrepancies which might grow up to become a problem for the SM
- It is hard to believe that the quest for the miracle of Nature is over

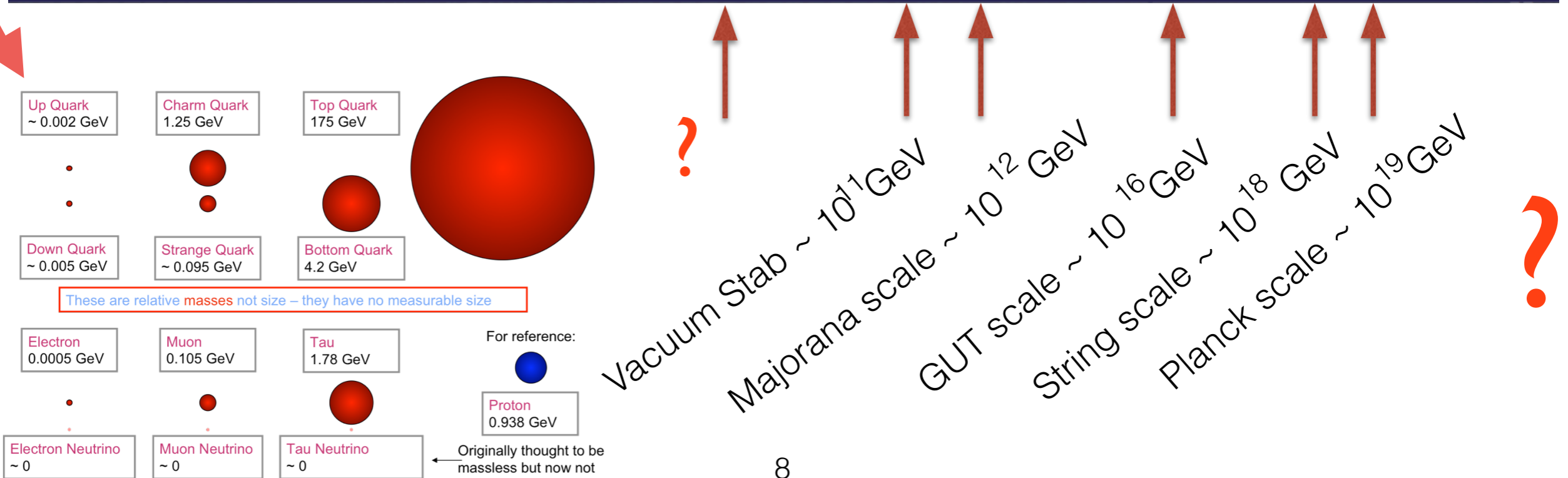
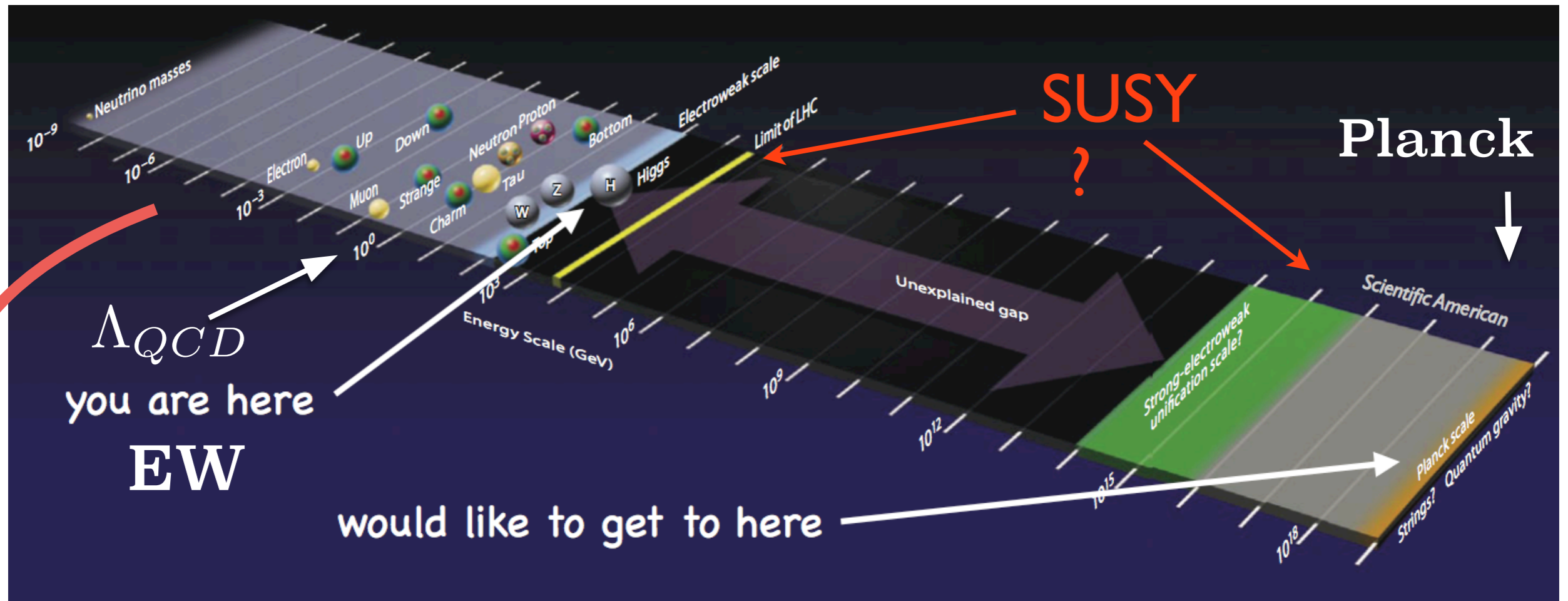
ACTUAL QUESTIONS

- What the dark matter is made of?
- Nature of neutrino: Dirac or Majorana ?
- The Higgs sector: one or many?
- CP violation and baryon asymmetry of the Universe?
- Is there and what kind of confinement-deconfinement phase transition?
- Lepton nonuniversality: fiction or real?
- How are hadrons build: spin, multi quark states, glueballs?
- (Non)stability of electroweak vacuum?
- Is there any deviations from the Standard model ?

Is there another scale except for EW and Planck?



Is there another scale except for EW and Planck?



THE LAGRANGIAN

$$\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}_{Yukawa} + \mathcal{L}_{Higgs},$$

$$\mathcal{L}_{gauge} = -\frac{1}{4}G_{\mu\nu}^a G_{\mu\nu}^a - \frac{1}{4}W_{\mu\nu}^i W_{\mu\nu}^i - \frac{1}{4}B_{\mu\nu} B_{\mu\nu}$$

$$+ i\bar{L}_\alpha \gamma^\mu D_\mu L_\alpha + i\bar{Q}_\alpha \gamma^\mu D_\mu Q_\alpha + i\bar{E}_\alpha \gamma^\mu D_\mu E_\alpha$$

$$+ i\bar{U}_\alpha \gamma^\mu D_\mu U_\alpha + i\bar{D}_\alpha \gamma^\mu D_\mu D_\alpha + (D_\mu H)^\dagger (D_\mu H),$$

$$+ i\bar{N}_\alpha \gamma^\mu \partial_\mu N_\alpha \quad \leftarrow \text{possible right handed neutrino ?}$$

$$\mathcal{L}_{Yukawa} = y_{\alpha\beta}^L \bar{L}_\alpha E_\beta H + y_{\alpha\beta}^D \bar{Q}_\alpha D_\beta H + y_{\alpha\beta}^U \bar{Q}_\alpha U_\beta \tilde{H} + h.c.,$$

$$+ y_{\alpha\beta}^N \bar{L}_\alpha N_\beta \tilde{H} \quad \leftarrow$$

$$\mathcal{L}_{Higgs} = -V = m^2 H^\dagger H - \frac{\lambda}{2} (H^\dagger H)^2$$

THE LAGRANGIAN

$$\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}$$

$$\mathcal{L}_{gauge} = -\frac{1}{4}G_{\mu\nu}^c G_{\mu\nu}^c$$

$$+ i\bar{L}_\alpha \gamma^\mu D_\mu L_\alpha$$

$$+ i\bar{U}_\alpha \gamma^\mu D_\mu U_\alpha +$$

$$+ i\bar{N}_\alpha \gamma^\mu \partial_\mu N_\alpha \quad \leftarrow$$

$$\mathcal{L}_{Yukawa} = y_{\alpha\beta}^L \bar{L}_\alpha E_\beta H$$

$$+ y_{\alpha\beta}^N \bar{L}_\alpha \Lambda_\beta$$

$$\mathcal{L}_{Higgs} = -V =$$

$$\begin{aligned} \mathcal{L}_{SM} = & -\frac{1}{2}\partial_\nu g_\mu^a \partial_\nu g_\mu^a - g_s f^{abc} \partial_\mu g_\nu^a g_\mu^b g_\nu^c - \frac{1}{4}g_s^2 f^{abc} f^{ade} g_\mu^b g_\nu^c g_\mu^d g_\nu^e - \partial_\nu W_\mu^+ \partial_\nu W_\mu^- - \\ & M^2 W_\mu^+ W_\mu^- - \frac{1}{2}\partial_\nu Z_\mu^0 \partial_\nu Z_\mu^0 - \frac{1}{2c_w^2} M^2 Z_\mu^0 Z_\mu^0 - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - igc_w (\partial_\nu Z_\mu^0 (W_\mu^+ W_\nu^- - \\ & W_\nu^+ W_\mu^-) - Z_\nu^0 (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + Z_\mu^0 (W_\nu^+ \partial_\mu W_\nu^- - W_\nu^- \partial_\mu W_\nu^+)) - \\ & ig s_w (\partial_\nu A_\mu (W_\mu^+ W_\nu^- - W_\nu^+ W_\mu^-) - A_\nu (W_\mu^+ \partial_\nu W_\mu^- - W_\mu^- \partial_\nu W_\mu^+) + A_\mu (W_\nu^+ \partial_\nu W_\mu^- - \\ & W_\nu^- \partial_\nu W_\mu^+)) - \frac{1}{2}g^2 W_\mu^+ W_\mu^- W_\nu^+ W_\nu^- + \frac{1}{2}g^2 W_\mu^+ W_\nu^- W_\mu^- W_\nu^+ + g^2 c_w^2 (Z_\mu^0 W_\mu^+ Z_\nu^0 W_\nu^- - \\ & Z_\mu^0 Z_\nu^0 W_\mu^+ W_\nu^-) + g^2 s_w^2 (A_\mu W_\mu^+ A_\nu W_\nu^- - A_\mu A_\nu W_\mu^+ W_\nu^-) + g^2 s_w c_w (A_\mu Z_\nu^0 (W_\mu^+ W_\nu^- - \\ & W_\nu^+ W_\mu^-) - 2A_\mu Z_\mu^0 W_\nu^+ W_\nu^-) - \frac{1}{2}\partial_\mu H \partial_\mu H - 2M^2 \alpha_h H^2 - \partial_\mu \phi^+ \partial_\mu \phi^- - \frac{1}{2}\partial_\mu \phi^0 \partial_\mu \phi^0 - \\ & \beta_h \left(\frac{2M^2}{g^2} + \frac{2M}{g} H + \frac{1}{2}(H^2 + \phi^0 \phi^0 + 2\phi^+ \phi^-) \right) + \frac{2M^4}{g^2} \alpha_h - \\ & g\alpha_h M (H^3 + H\phi^0 \phi^0 + 2H\phi^+ \phi^-) - \\ & \frac{1}{8}g^2 \alpha_h (H^4 + (\phi^0)^4 + 4(\phi^+ \phi^-)^2 + 4(\phi^0)^2 \phi^+ \phi^- + 4H^2 \phi^+ \phi^- + 2(\phi^0)^2 H^2) - \\ & gM W_\mu^+ W_\mu^- H - \frac{1}{2}g \frac{M}{c_w^2} Z_\mu^0 Z_\mu^0 H - \\ & \frac{1}{2}ig (W_\mu^+ (\phi^0 \partial_\mu \phi^- - \phi^- \partial_\mu \phi^0) - W_\mu^- (\phi^0 \partial_\mu \phi^+ - \phi^+ \partial_\mu \phi^0)) + \\ & \frac{1}{2}g (W_\mu^+ (H \partial_\mu \phi^- - \phi^- \partial_\mu H) + W_\mu^- (H \partial_\mu \phi^+ - \phi^+ \partial_\mu H)) + \frac{1}{2}g \frac{1}{c_w} (Z_\mu^0 (H \partial_\mu \phi^0 - \phi^0 \partial_\mu H) + \\ & M (\frac{1}{c_w} Z_\mu^0 \partial_\mu \phi^0 + W_\mu^+ \partial_\mu \phi^- + W_\mu^- \partial_\mu \phi^+)) - ig \frac{s_w^2}{c_w} M Z_\mu^0 (W_\mu^+ \phi^- - W_\mu^- \phi^+) + ig s_w M A_\mu (W_\mu^+ \phi^- - \\ & W_\mu^- \phi^+) - ig \frac{1-2c_w^2}{2c_w} Z_\mu^0 (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) + ig s_w A_\mu (\phi^+ \partial_\mu \phi^- - \phi^- \partial_\mu \phi^+) - \\ & \frac{1}{4}g^2 W_\mu^+ W_\mu^- (H^2 + (\phi^0)^2 + 2\phi^+ \phi^-) - \frac{1}{8}g^2 \frac{1}{c_w^2} Z_\mu^0 Z_\mu^0 (H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)\phi^+ \phi^-) - \\ & \frac{1}{2}g^2 \frac{s_w^2}{c_w} Z_\mu^0 \phi^0 (W_\mu^+ \phi^- + W_\mu^- \phi^+) - \frac{1}{2}ig^2 \frac{s_w^2}{c_w} Z_\mu^0 H (W_\mu^+ \phi^- - W_\mu^- \phi^+) + \frac{1}{2}g^2 s_w A_\mu \phi^0 (W_\mu^+ \phi^- + \\ & W_\mu^- \phi^+) + \frac{1}{2}ig^2 s_w A_\mu H (W_\mu^+ \phi^- - W_\mu^- \phi^+) - g^2 \frac{s_w}{c_w} (2c_w^2 - 1) Z_\mu^0 A_\mu \phi^+ \phi^- - \\ & g^2 s_w^2 A_\mu A_\mu \phi^+ \phi^- + \frac{1}{2}ig s_w \lambda_{ij}^a (\bar{q}_i^\sigma \gamma^\mu q_j^\sigma) g_\mu^a - \bar{e}^\lambda (\gamma^\partial + m_e^\lambda) e^\lambda - \bar{\nu}^\lambda (\gamma^\partial + m_\nu^\lambda) \nu^\lambda - \bar{u}_j^\lambda (\gamma^\partial + \\ & m_u^\lambda) u_j^\lambda - \bar{d}_j^\lambda (\gamma^\partial + m_d^\lambda) d_j^\lambda + ig s_w A_\mu (-\bar{e}^\lambda \gamma^\mu e^\lambda) + \frac{2}{3}(\bar{u}_j^\lambda \gamma^\mu u_j^\lambda) - \frac{1}{3}(\bar{d}_j^\lambda \gamma^\mu d_j^\lambda) + \\ & \frac{ig}{4c_w} Z_\mu^0 \{ (\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{e}^\lambda \gamma^\mu (4s_w^2 - 1 - \gamma^5) e^\lambda) + (\bar{d}_j^\lambda \gamma^\mu (\frac{4}{3}s_w^2 - 1 - \gamma^5) d_j^\lambda) + \\ & (\bar{u}_j^\lambda \gamma^\mu (1 - \frac{8}{3}s_w^2 + \gamma^5) u_j^\lambda) \} + \frac{ig}{2\sqrt{2}} W_\mu^+ ((\bar{\nu}^\lambda \gamma^\mu (1 + \gamma^5) U^{lep}{}_{\lambda\kappa} e^\kappa) + (\bar{u}_j^\lambda \gamma^\mu (1 + \gamma^5) C_{\lambda\kappa} d_j^\kappa)) + \\ & \frac{ig}{2\sqrt{2}} W_\mu^- ((\bar{e}^\kappa U^{lep}{}_{\kappa\lambda}^\dagger \gamma^\mu (1 + \gamma^5) \nu^\lambda) + (\bar{d}_j^\kappa C_{\kappa\lambda}^\dagger \gamma^\mu (1 + \gamma^5) u_j^\lambda)) + \\ & \frac{ig}{2M\sqrt{2}} \phi^+ (-m_e^\kappa (\bar{\nu}^\lambda U^{lep}{}_{\lambda\kappa} (1 - \gamma^5) e^\kappa) + m_\nu^\lambda (\bar{\nu}^\lambda U^{lep}{}_{\lambda\kappa} (1 + \gamma^5) e^\kappa) + \\ & \frac{ig}{2M\sqrt{2}} \phi^- (m_e^\lambda (\bar{e}^\lambda U^{lep}{}_{\lambda\kappa}^\dagger (1 + \gamma^5) \nu^\kappa) - m_\nu^\kappa (\bar{e}^\lambda U^{lep}{}_{\lambda\kappa}^\dagger (1 - \gamma^5) \nu^\kappa) - \frac{g}{2} \frac{m_\nu^\lambda}{M} H (\bar{\nu}^\lambda \nu^\lambda) - \\ & \frac{g}{2} \frac{m_d^\lambda}{M} H (\bar{e}^\lambda e^\lambda) + \frac{ig}{2} \frac{m_\nu^\lambda}{M} \phi^0 (\bar{\nu}^\lambda \gamma^5 \nu^\lambda) - \frac{ig}{2} \frac{m_e^\lambda}{M} \phi^0 (\bar{e}^\lambda \gamma^5 e^\lambda) - \frac{1}{4} \bar{\nu}_\lambda M_{\lambda\kappa}^R (1 - \gamma^5) \hat{\nu}_\kappa - \\ & \frac{1}{4} \bar{\nu}_\lambda M_{\lambda\kappa}^R (1 - \gamma^5) \hat{\nu}_\kappa + \frac{ig}{2M\sqrt{2}} \phi^+ (-m_d^\kappa (\bar{u}_j^\lambda C_{\lambda\kappa} (1 - \gamma^5) d_j^\kappa) + m_u^\lambda (\bar{u}_j^\lambda C_{\lambda\kappa} (1 + \gamma^5) d_j^\kappa) + \\ & \frac{ig}{2M\sqrt{2}} \phi^- (m_d^\lambda (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 + \gamma^5) u_j^\kappa) - m_u^\kappa (\bar{d}_j^\lambda C_{\lambda\kappa}^\dagger (1 - \gamma^5) u_j^\kappa) - \frac{g}{2} \frac{m_u^\lambda}{M} H (\bar{u}_j^\lambda u_j^\lambda) - \\ & \frac{g}{2} \frac{m_d^\lambda}{M} H (\bar{d}_j^\lambda d_j^\lambda) + \frac{ig}{2} \frac{m_u^\lambda}{M} \phi^0 (\bar{u}_j^\lambda \gamma^5 u_j^\lambda) - \frac{ig}{2} \frac{m_d^\lambda}{M} \phi^0 (\bar{d}_j^\lambda \gamma^5 d_j^\lambda) + \bar{G}^a \partial^2 G^a + g_s f^{abc} \partial_\mu \bar{G}^a G^b g_\mu^c + \\ & \bar{X}^+ (\partial^2 - M^2) X^+ + \bar{X}^- (\partial^2 - M^2) X^- + \bar{X}^0 (\partial^2 - \frac{M^2}{c_w^2}) X^0 + \bar{Y} \partial^2 Y + igc_w W_\mu^+ (\partial_\mu \bar{X}^0 X^- - \\ & \partial_\mu \bar{X}^+ X^0) + ig s_w W_\mu^+ (\partial_\mu \bar{Y} X^- - \partial_\mu \bar{X}^+ Y) + igc_w W_\mu^- (\partial_\mu \bar{X}^- X^0 - \\ & \partial_\mu \bar{X}^0 X^+) + ig s_w W_\mu^- (\partial_\mu \bar{X}^- Y - \partial_\mu \bar{Y} X^+) + igc_w Z_\mu^0 (\partial_\mu \bar{X}^+ X^+ - \\ & \partial_\mu \bar{X}^- X^-) + ig s_w A_\mu (\partial_\mu \bar{X}^+ X^+ - \\ & \partial_\mu \bar{X}^- X^-) - \frac{1}{2}gM (\bar{X}^+ X^+ H + \bar{X}^- X^- H + \frac{1}{c_w^2} \bar{X}^0 X^0 H) + \frac{1-2c_w^2}{2c_w} igM (\bar{X}^+ X^0 \phi^+ - \bar{X}^- X^0 \phi^-) + \\ & \frac{1}{2c_w} igM (\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-) + igM s_w (\bar{X}^0 X^- \phi^+ - \bar{X}^0 X^+ \phi^-) + \\ & \frac{1}{2}igM (\bar{X}^+ X^+ \phi^0 - \bar{X}^- X^- \phi^0) . \end{aligned}$$

THE LAGRANGIAN

$$\mathcal{L} = \mathcal{L}_{gauge} + \mathcal{L}_{Yukawa} + \mathcal{L}_{Higgs},$$

$$\mathcal{L}_{gauge} = -\frac{1}{4}G_{\mu\nu}^a G_{\mu\nu}^a - \frac{1}{4}W_{\mu\nu}^i W_{\mu\nu}^i - \frac{1}{4}B_{\mu\nu} B_{\mu\nu}$$

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Three gauge couplings

THE LAGRANGIAN

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Three gauge couplings

Three or four Yukawa matrices

THE LAGRANGIAN

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Three gauge couplings

Three or four Yukawa matrices

Two parameters

THE LAGRANGIAN

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possible right handed neutrino ?

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$$\mathcal{L}_{Higgs} = -V = m^2 H^\dagger H - \frac{\lambda}{2} (H^\dagger H)^2$$

All these parameters are not predicted by the SM and determined experimentally

Three gauge couplings

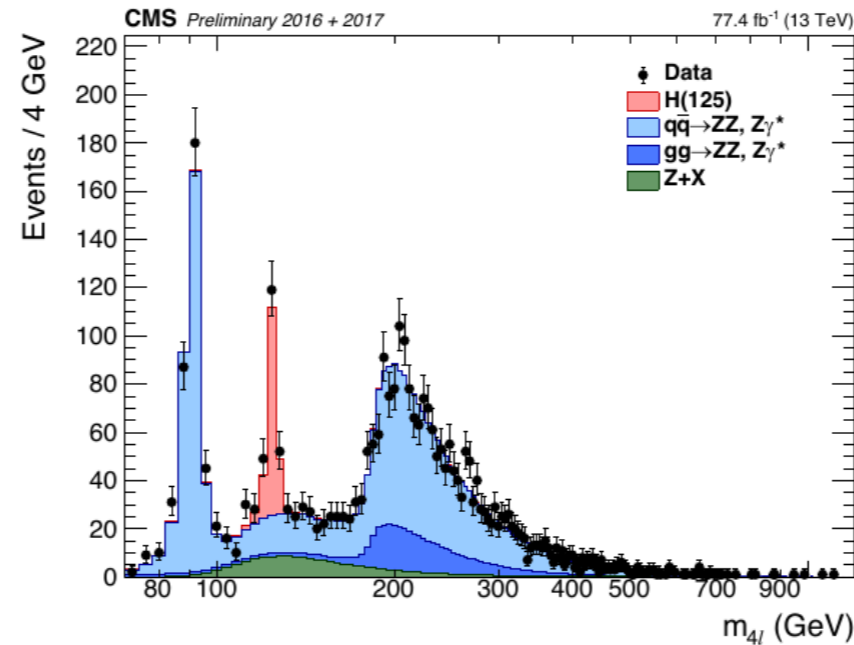
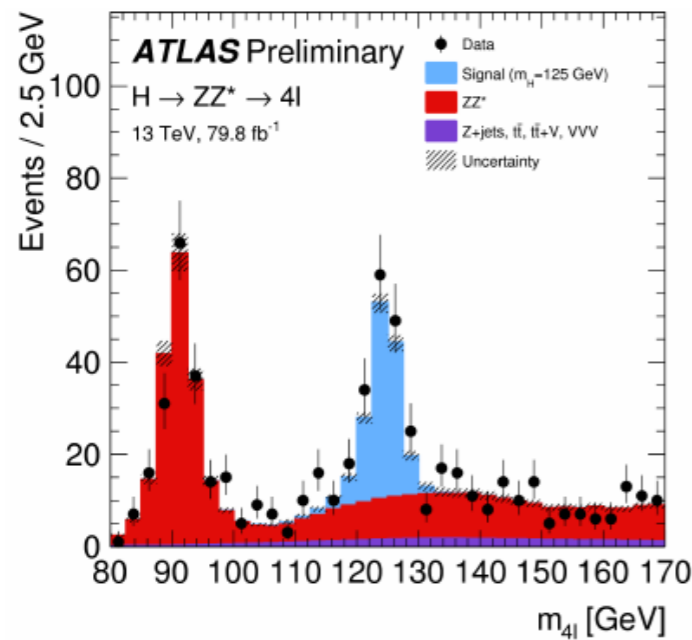
Three or four Yukawa matrices

Two parameters

Higgs sector

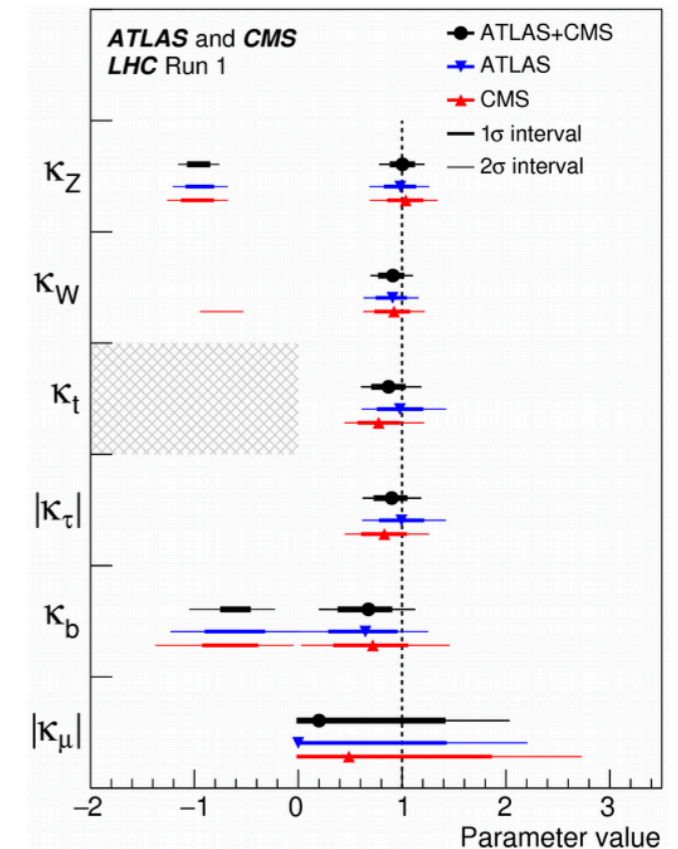
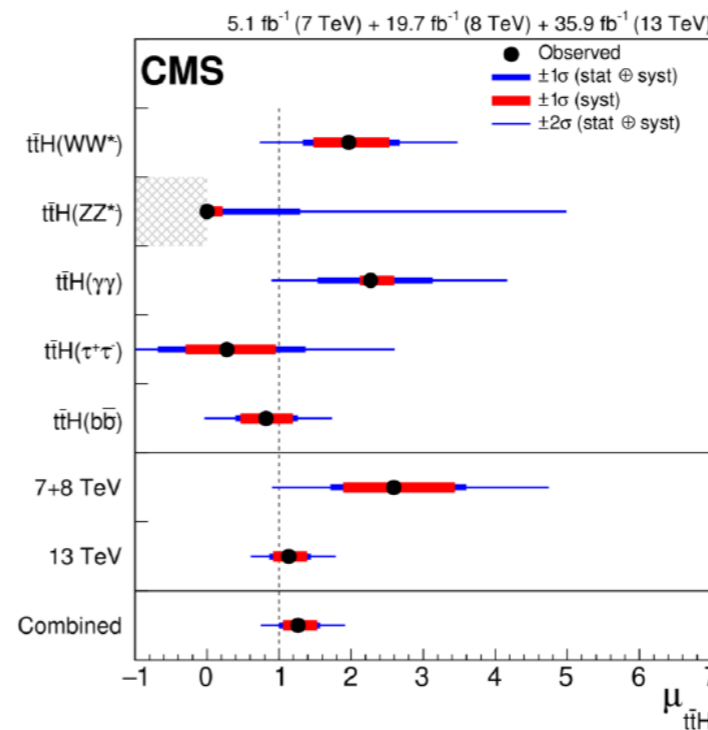
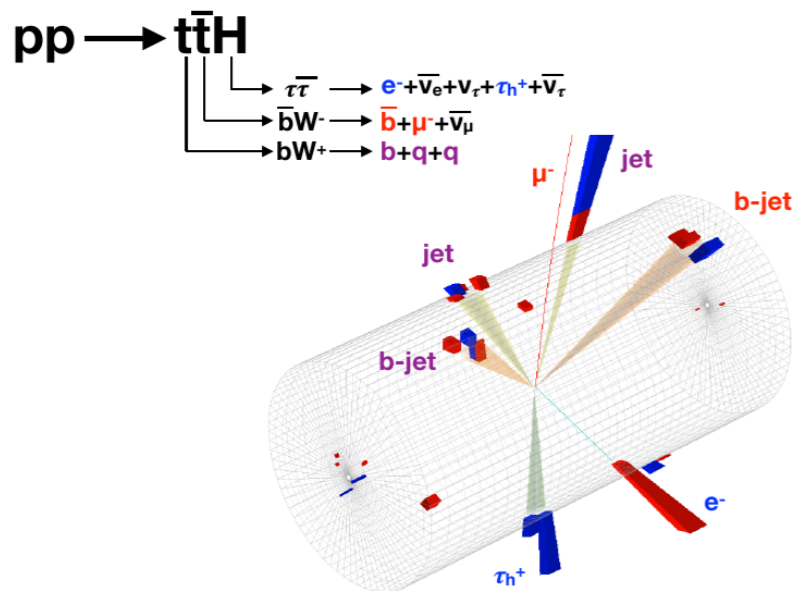
Higgs bosons - entering precision era

Run-2 analyses with 80 fb^{-1} for the first time – higher precision is coming!



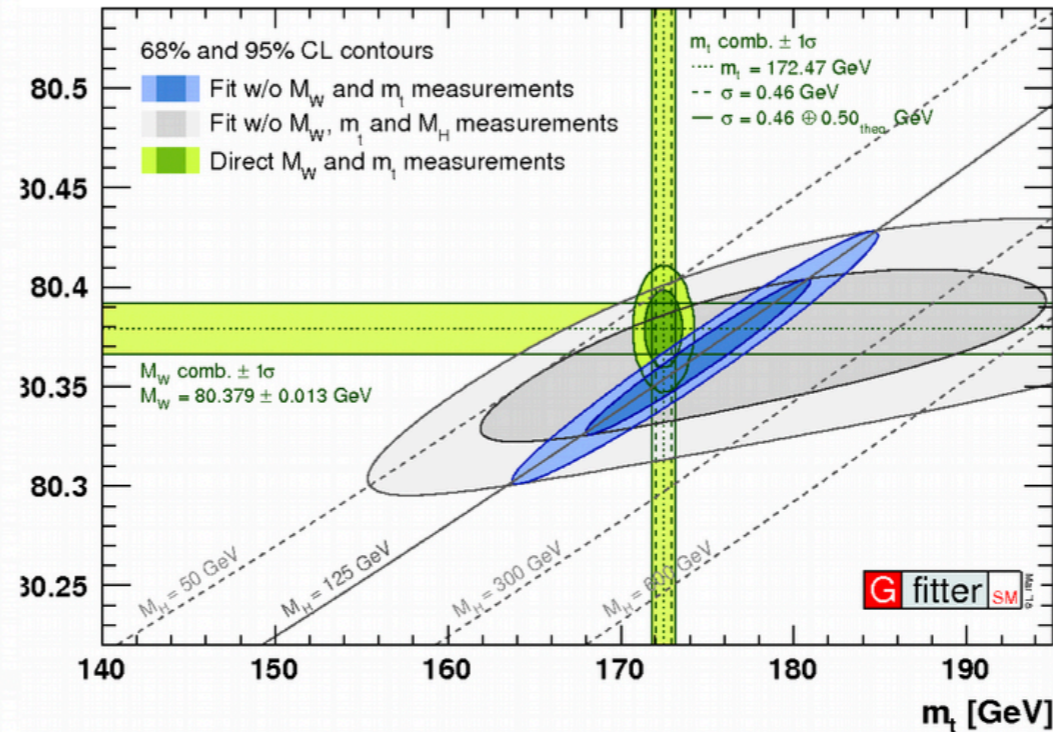
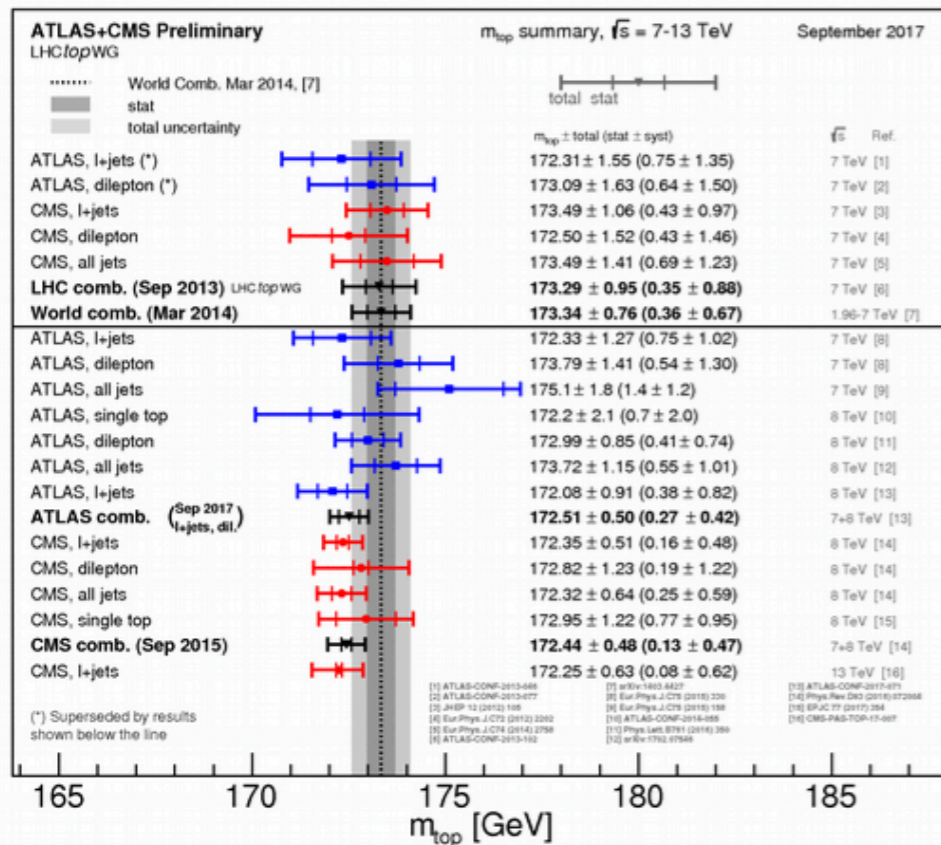
D. Charlton
LHCp2018

ttH observation



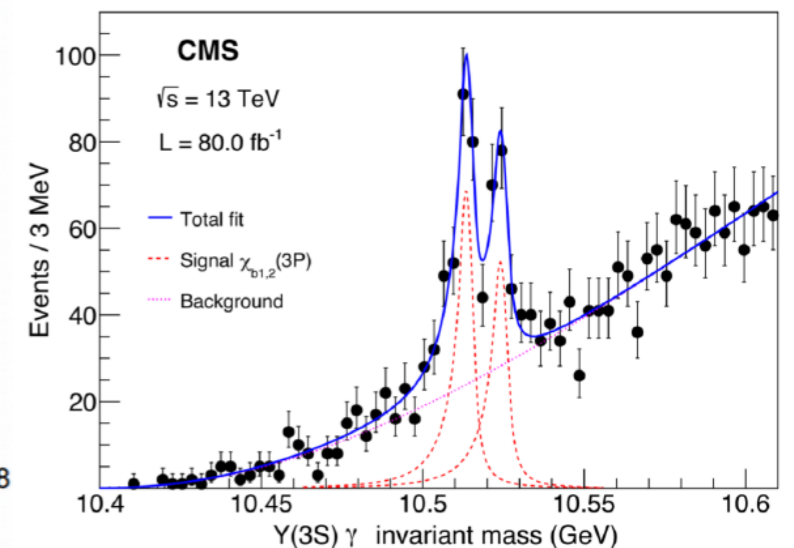
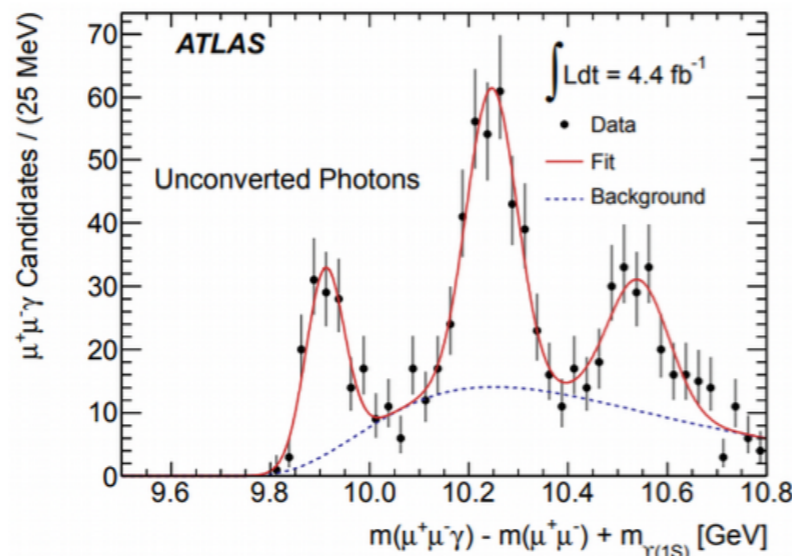
Precision EW mass measurements

D. Charlton
LHCp2018



Precision spectroscopy!

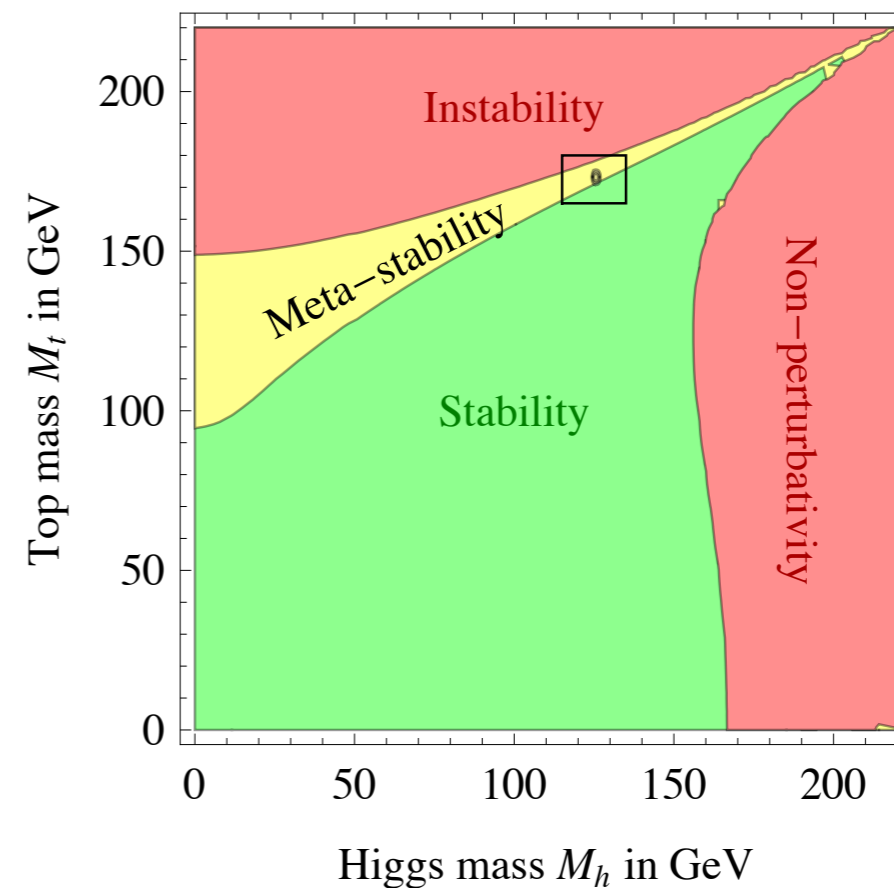
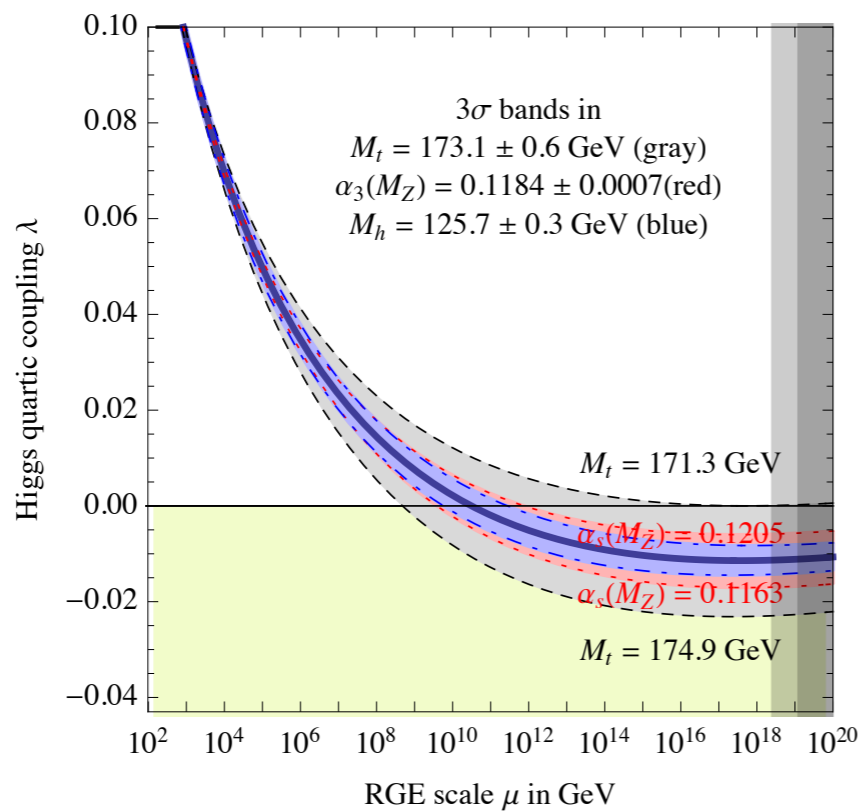
$$m(\chi_{b2}(3P)) - m(\chi_{b2}(3P)) = 10.60 \pm 0.64(\text{stat}) \pm 0.17(\text{syst}) \text{ MeV}$$



THE STANDARD MODEL: THE STATUS REPORT AND OPEN QUESTIONS

📌 The electroweak vacuum is unstable under radiative corrections

📌 The whole construction of the SM may be in trouble being metastable or even unstable



📌 the situation crucially depends on the top and Higgs mass values and requires severe fine-tuning and high accuracy of calculations (3 loops)

g-2

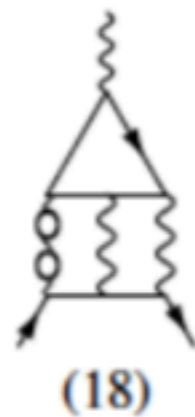
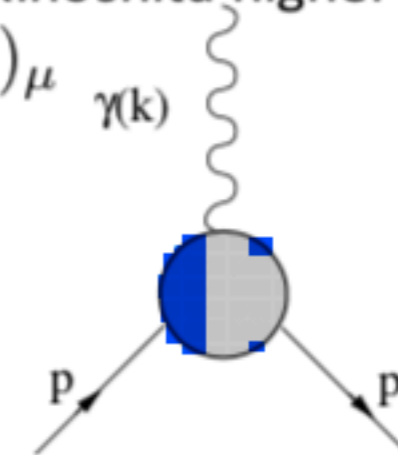
Muon anomalous magnetic moment

$$ie\bar{u}_\ell(p') \left[\gamma^\mu - \frac{a_\ell}{2m_\ell} i\sigma^{\mu\nu} q_\nu \right] u_\ell(p) \epsilon_\mu^*, \quad q_\mu = (p - p')_\mu$$

Dirac equation predicts $g=2$ $a = (g - 2)/2$

For electron a_e theory and experiment agrees!

(Schwinger α/π ,
Kinoshita higher orders in α)



$$a_\mu^{th} - a_\mu^{exp} = -(3.06 \pm 0.76) \times 10^{-8} \quad 3,7 \sigma$$

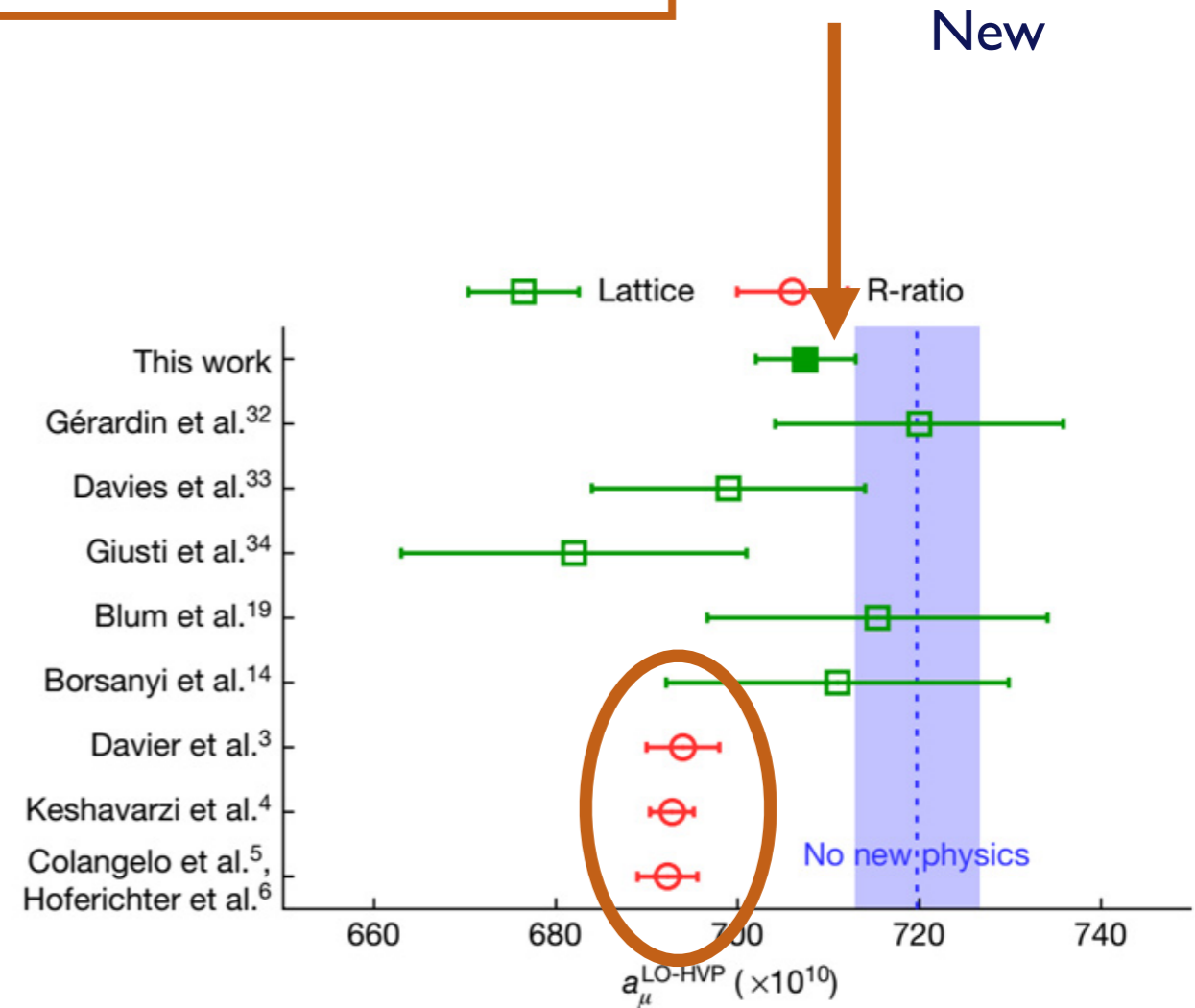
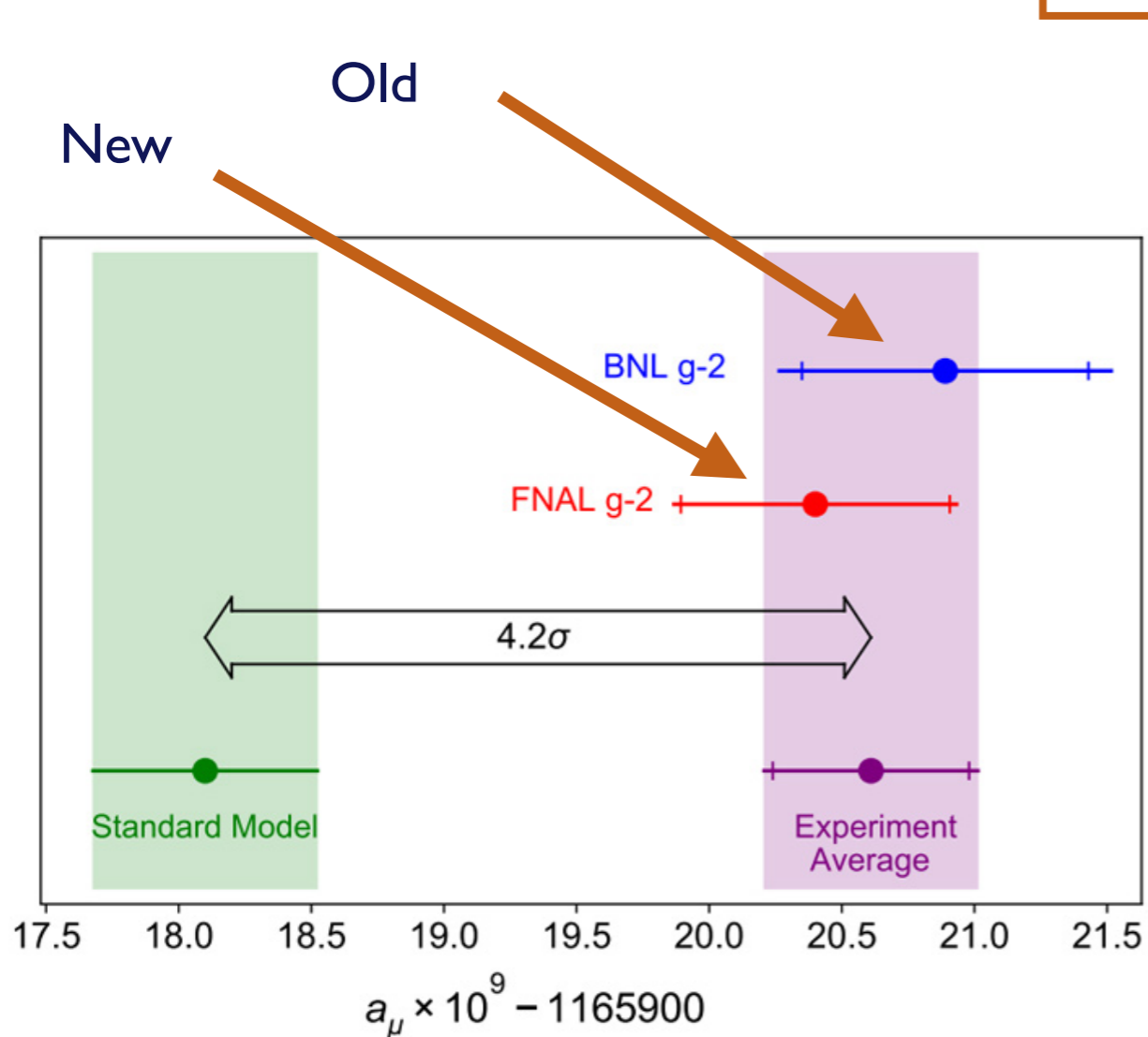
Theory: uncertainty in hadronic contributions to the muon $g - 2$, (Jägerlehner, 1802.08019).
Lattice QCD great progress light-by-light study (RBC & UKQCD, 1801.07224).

Fermilab and J-Park experiments are expected to clarify existing discrepancy!

THE STANDARD MODEL: THE STATUS AND OPEN QUESTIONS

FermiLab April 2021

$$a_\mu(\text{E821}) = (116\,592\,089 \pm 63) \times 10^{-11}.$$



The problem remained

May be not!

Conclusion: it is important to take into account the strong interactions contribution correctly!

$\Lambda \approx \nu$: SUSY and the muon ($g - 2$)

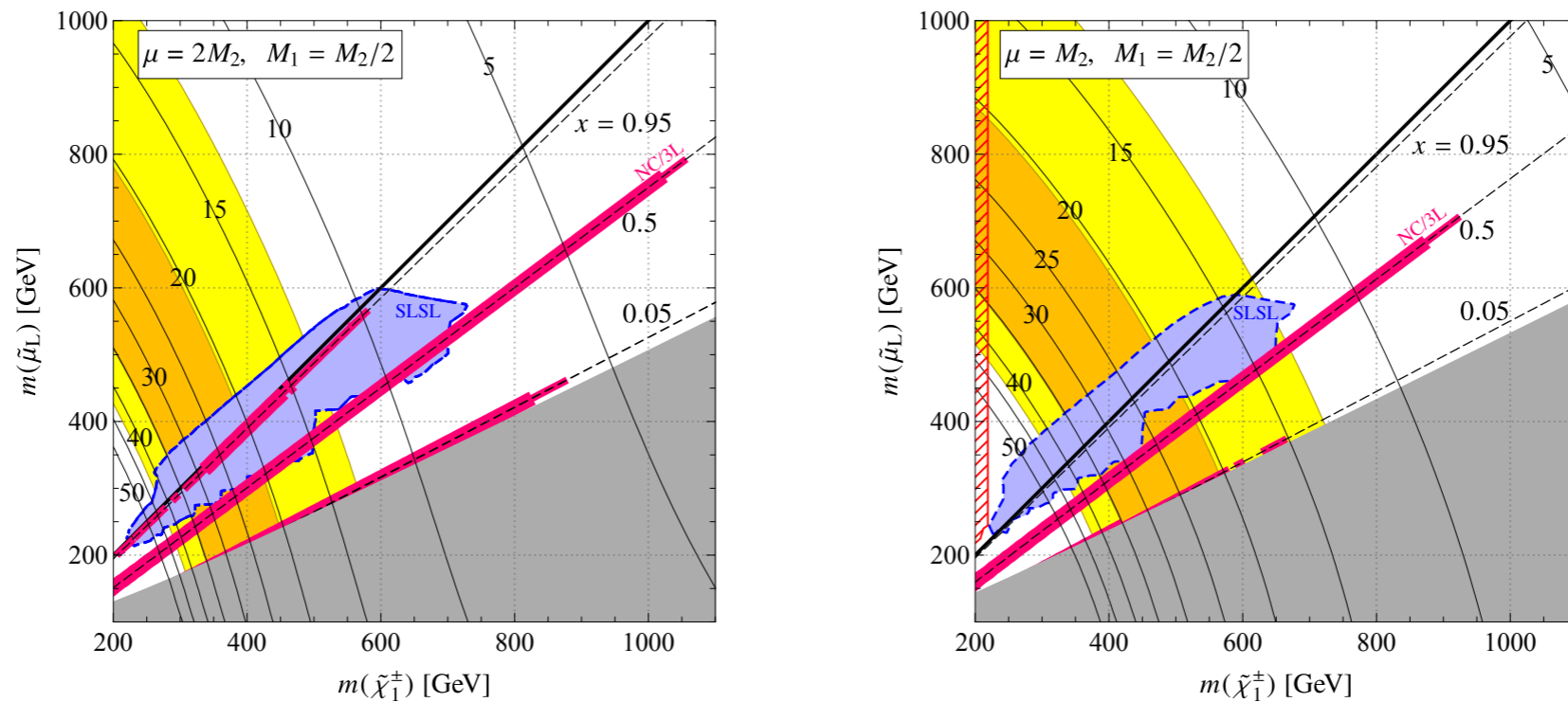


Figure: LHC Run 2 bounds on SUSY scenario for the muon $g - 2$ anomaly for $\tan \beta = 40$. Orange (yellow) regions satisfy the muon $g - 2$ anomaly at the 1σ (2σ) level [Endo et al., '20].

$(a_\mu^{\text{SM}})_{\text{weak}} \approx \frac{g^2 m_\mu^2}{32\pi^2 M_W^2} \approx 2 \times 10^{-9}$

$$a_\mu^{\text{SUSY}} \approx \frac{g^2 m_\mu^2 \tan \beta}{32\pi^2 \tilde{m}^2} \approx \underbrace{2 \times 10^{-9}}_{\tilde{m} = 500\text{GeV} \ \& \ \tan \beta = 40}$$

$\Lambda \lesssim 1$ GeV: Axion-like Particles and the muon ($g - 2$)

Axion-like Particle effective Lagrangian

$$\mathcal{L} = e^2 C_{\gamma\gamma} \frac{a}{\Lambda} F_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{c_{\mu\mu}}{2} \frac{\partial^\nu a}{\Lambda} \bar{\mu} \gamma_\nu \gamma_5 \mu$$

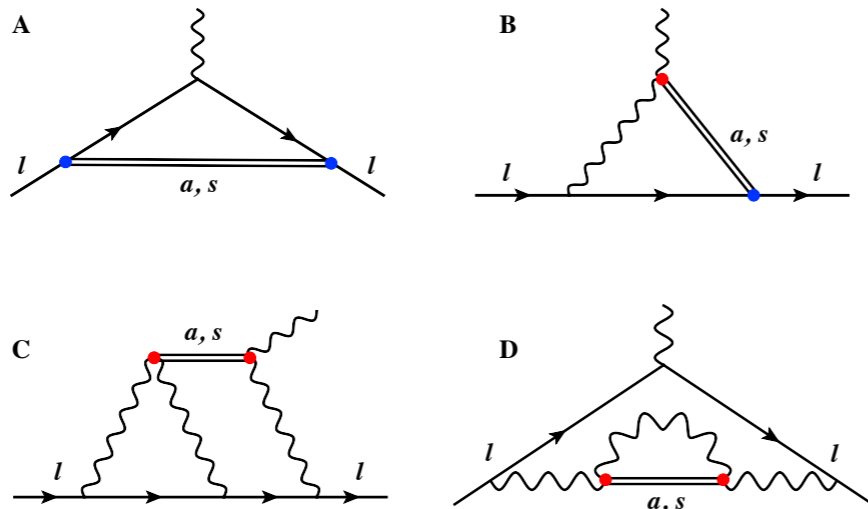


Figure: Contributions of a scalar ‘s’ and a pseudoscalar ‘a’ ALP to the $(g - 2)_\ell$.

[Marciano, Masiero, PP, Passera '16]

[Cornella, P.P., Sumensari '19]

$$\Delta a_\mu = \frac{m_\mu^2}{\Lambda^2} \left[\frac{12\alpha^3}{\pi} C_{\gamma\gamma}^2 \ln^2 \frac{\Lambda^2}{m_\mu^2} - \frac{(c_{\mu\mu})^2}{16\pi^2} h_1 \left(\frac{m_a^2}{m_\mu^2} \right) - \frac{2\alpha}{\pi} c_{\mu\mu} C_{\gamma\gamma} \ln \frac{\Lambda^2}{m_\mu^2} \right]$$

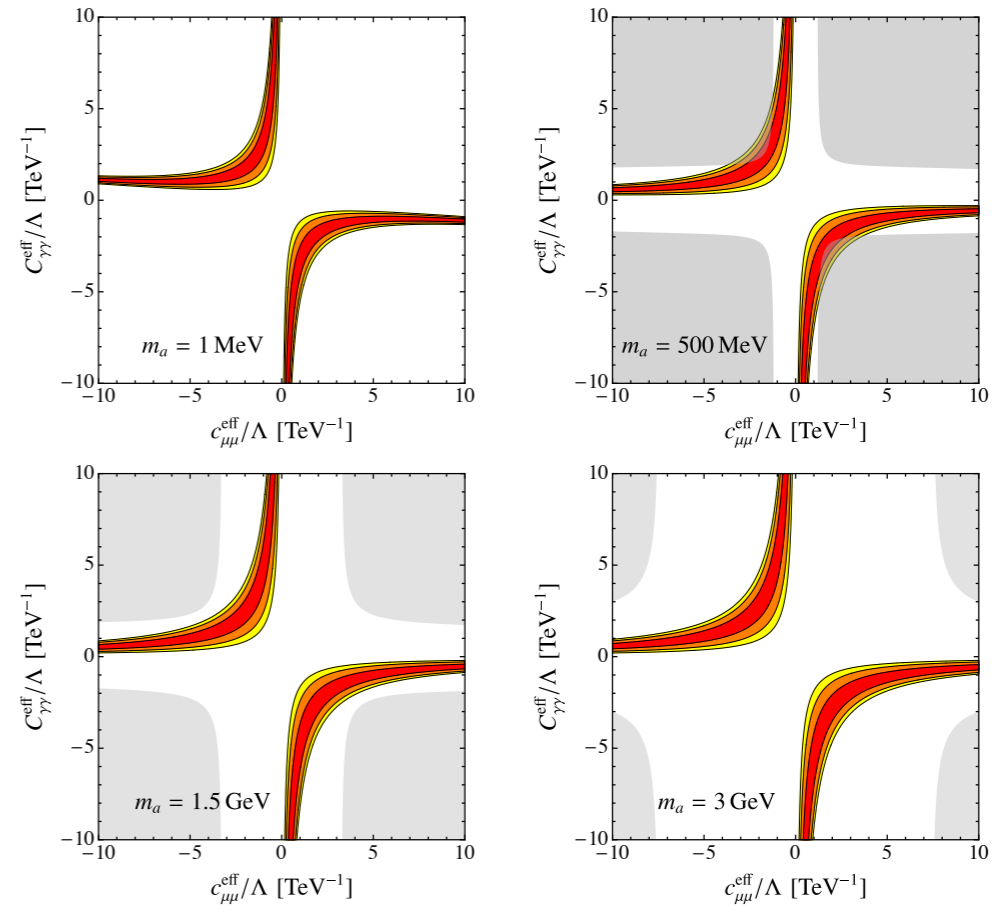
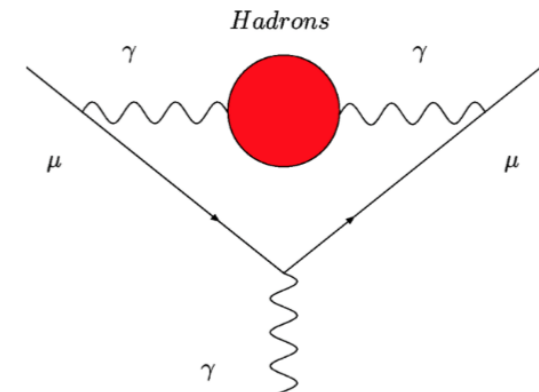


Figure: Δa_μ regions favoured at 68% (red), 95% (orange) and 99% (yellow) CL. Gray regions are excluded by the BaBar search $e^+e^- \rightarrow \mu^+\mu^- + \mu^+\mu^-$ [Bauer, Neubert, Thamm, '17]

HLO contribution from $e^+e^- \rightarrow \text{hadrons}$

- dominated by $e^+e^- \rightarrow \pi^+\pi^-$ channel (70% of the full hadronic)

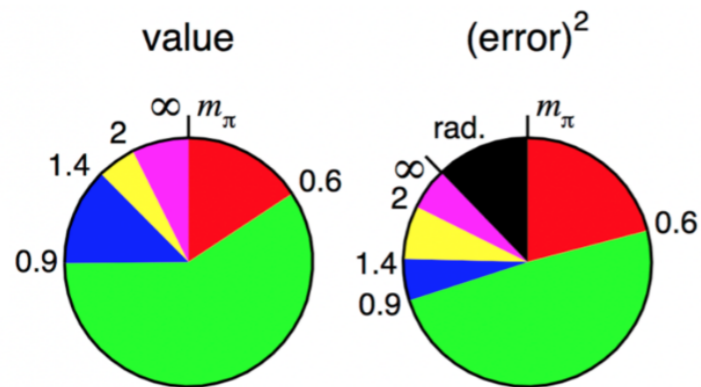


$$(a_\mu^{\text{HVP}})_{e^+e^-} = \frac{\alpha}{\pi^2} \int_{m_{\pi^0}^2}^{\infty} \frac{ds}{s} K(s) \text{Im} \Pi_{\text{had}}(s) = \frac{1}{4\pi^3} \int_{m_{\pi^0}^2}^{\infty} ds K(s) \sigma_{\text{had}}(s)$$

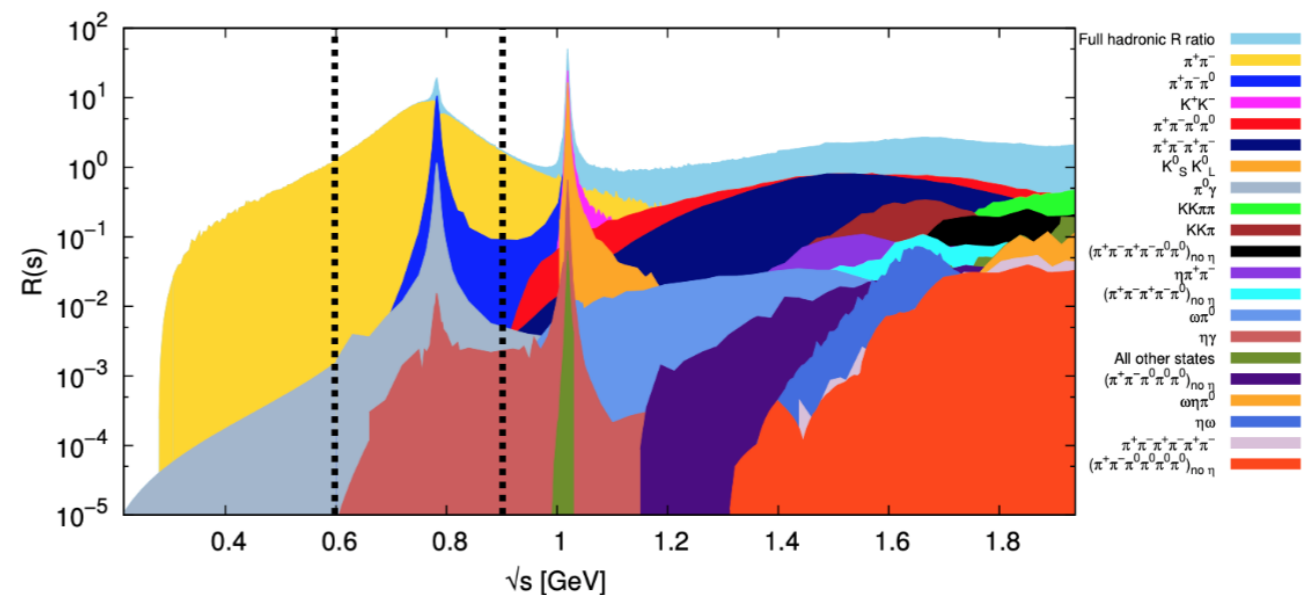
↓ dispersion relations
↓ optical theorem

kernel function
 $K(s) \approx m_\mu^2/3s$ for $\sqrt{s} \gg m_\mu$

$$\text{Im} \left[\text{wavy line} \text{---} \text{red circle} \text{---} \text{wavy line} \right] \sim \left| \text{wavy line} \text{---} \text{red lines} \right|^2 \sim \sigma(e^+e^- \rightarrow \gamma^* \rightarrow \text{hadrons})$$



Keshavarzi, Nomura, Teubner 2018

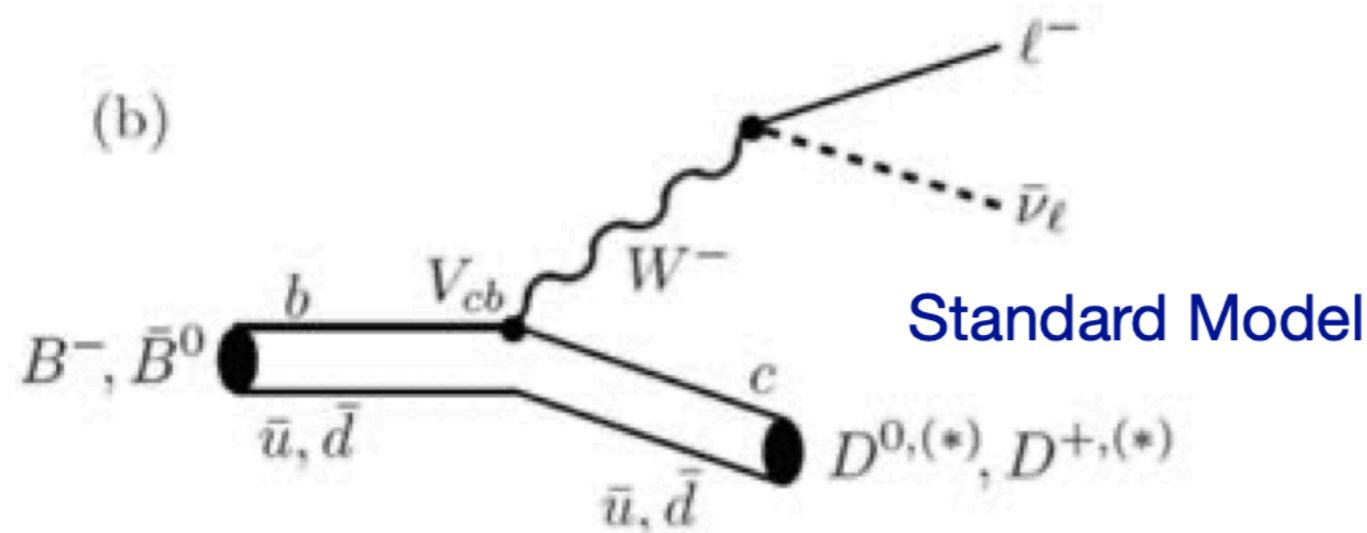


$$a_{\mu, e^+e^-}^{\text{HLO}} = 6931(40) \times 10^{-11} (0.6\%) \text{ [WP20]}$$

Flavor Anomaly

LEPTON (NON) UNIVERSALITY (!?)

Charged currents at tree level

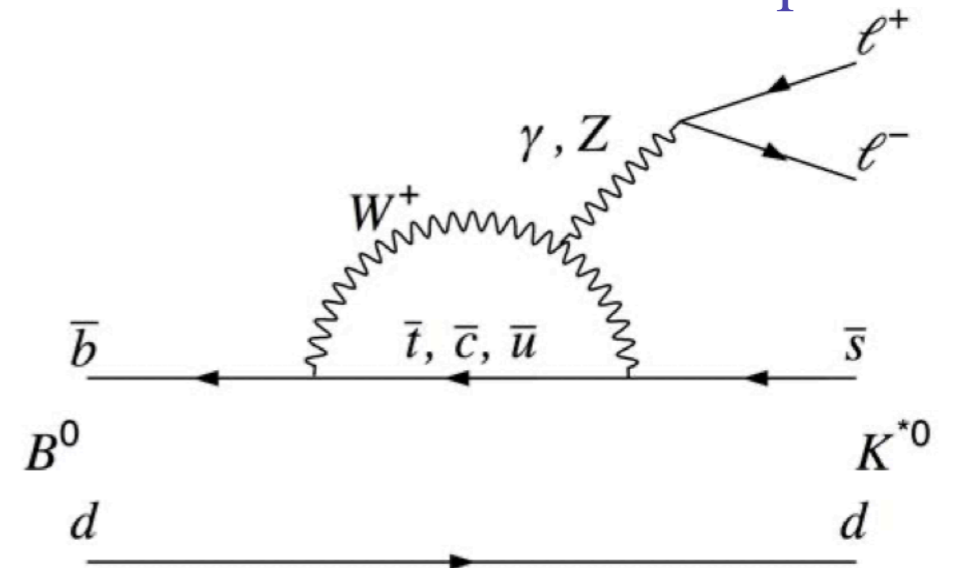


$$R_{D^{(*)}} = \frac{BR(B \rightarrow D^{(*)} \tau \nu_\tau)}{BR(B \rightarrow D^{(*)} \mu \nu_\mu)} \quad 3.8\sigma$$

$$\mathcal{R}(D)_{SM} = 0.299 \pm 0.003$$

$$\mathcal{R}(D^*)_{SM} = 0.258 \pm 0.005$$

Neutral currents at one-loop level



$$R_{K^{(*)}} = \frac{BR(B \rightarrow K^{(*)} \mu \mu)}{BR(B \rightarrow K^{(*)} e e)} \quad 2.5\sigma$$

$$\mathcal{R}(K^*)_{SM} = 1.0$$

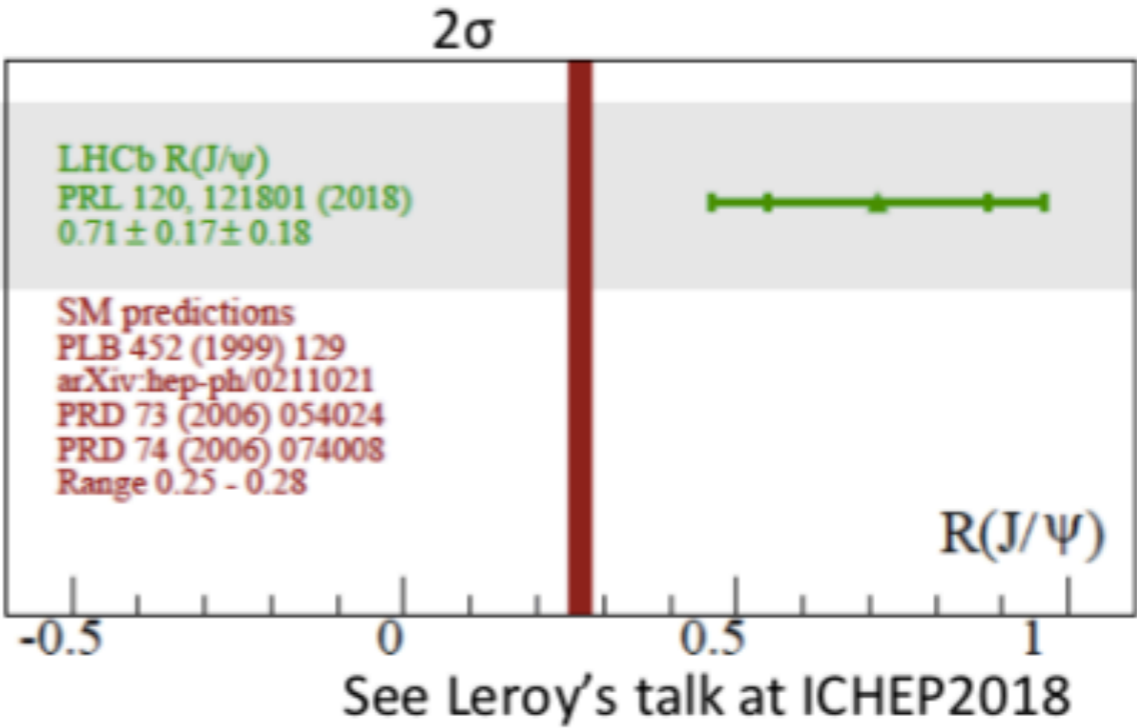
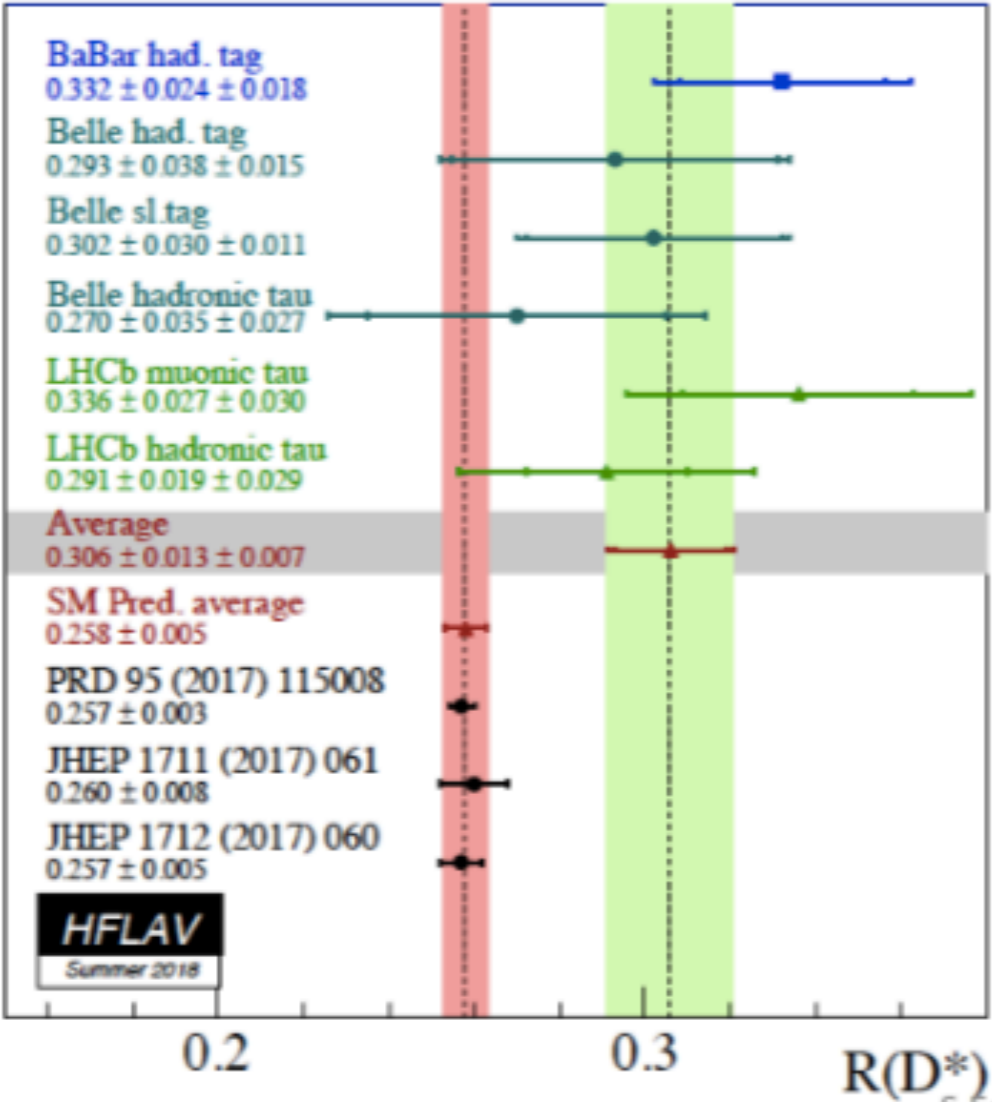
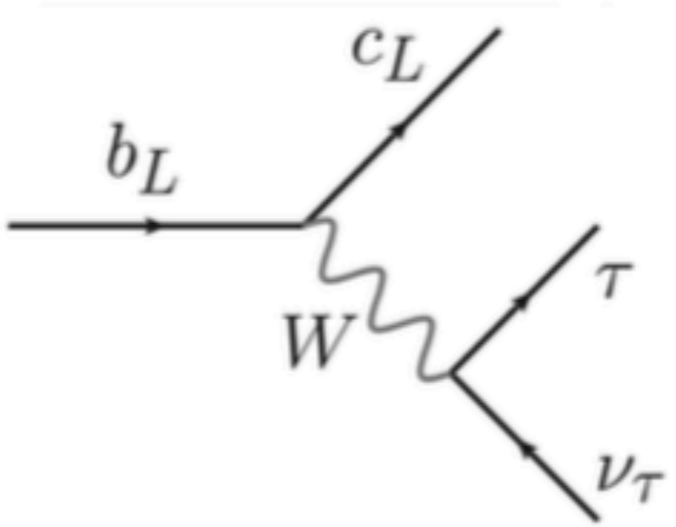
THE STANDARD MODEL: THE STATUS REPORT AND OPEN QUESTIONS

B physics anomalies: experimental results \neq SM predictions!

charged current (SM tree level)

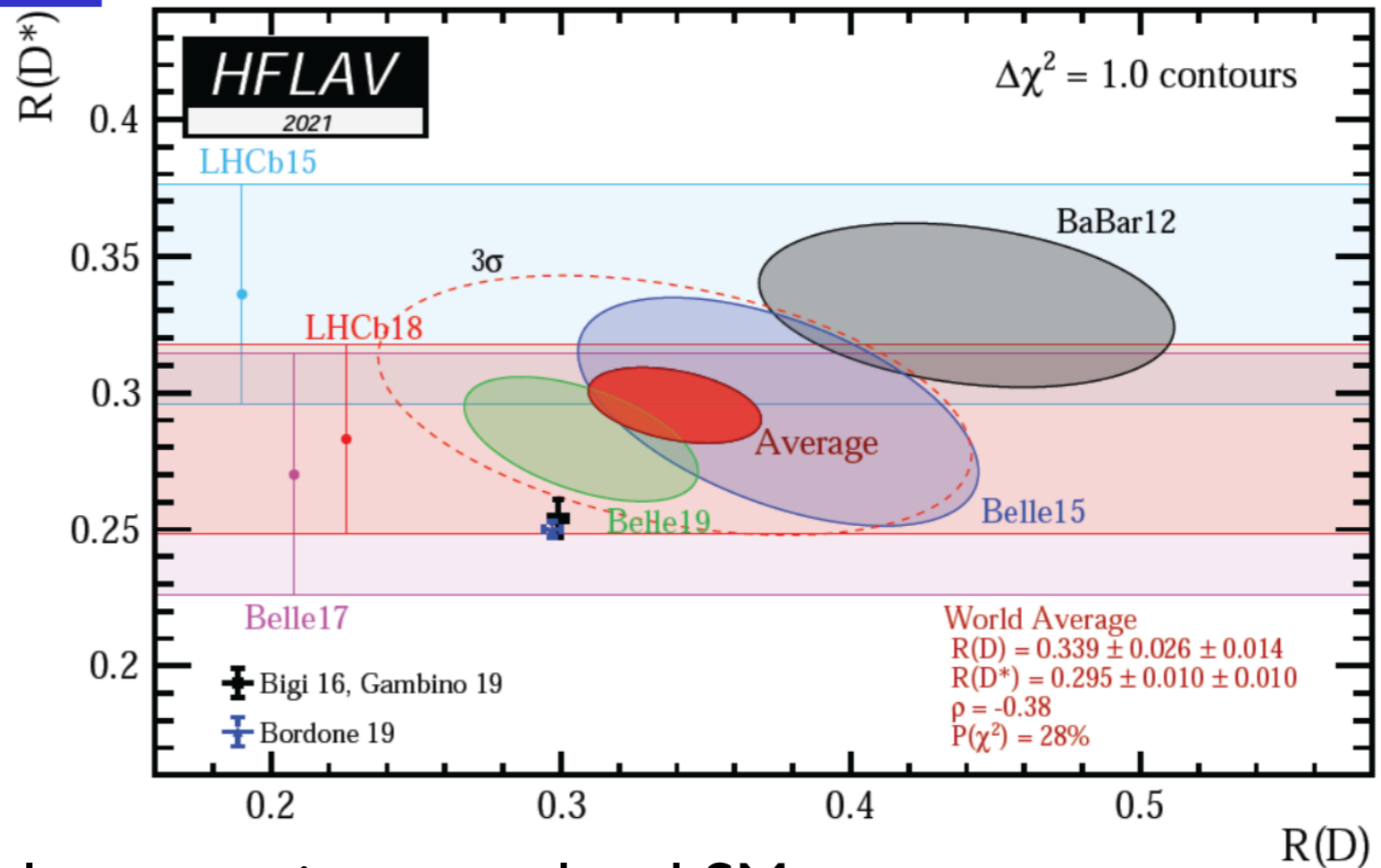
$$R_{D^{(*)}} = \frac{BR(B \rightarrow D^{(*)} \tau \nu_\tau)}{BR(B \rightarrow D^{(*)} \mu \nu_\mu)}$$

3.8 σ



$R(D)$ vs $R(D^*)$

- All experiments see an excess wrt SM predictions
- Combining $R(D)/R(D^*)$ average $\sim 3.4 \sigma$ tension with SM



- Intriguing as anomaly occurs in a tree-level SM process

- New LHCb result

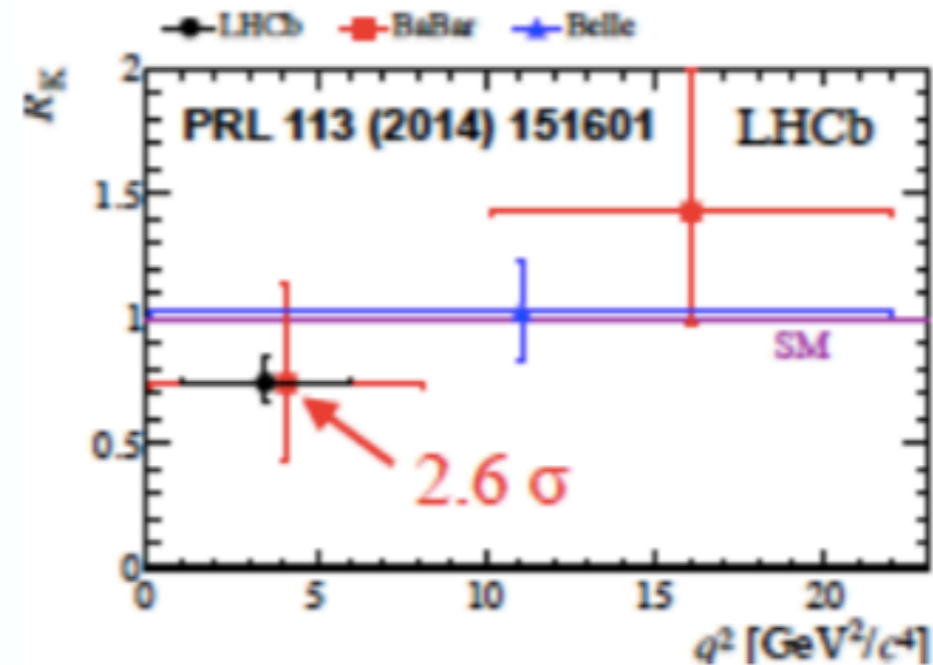
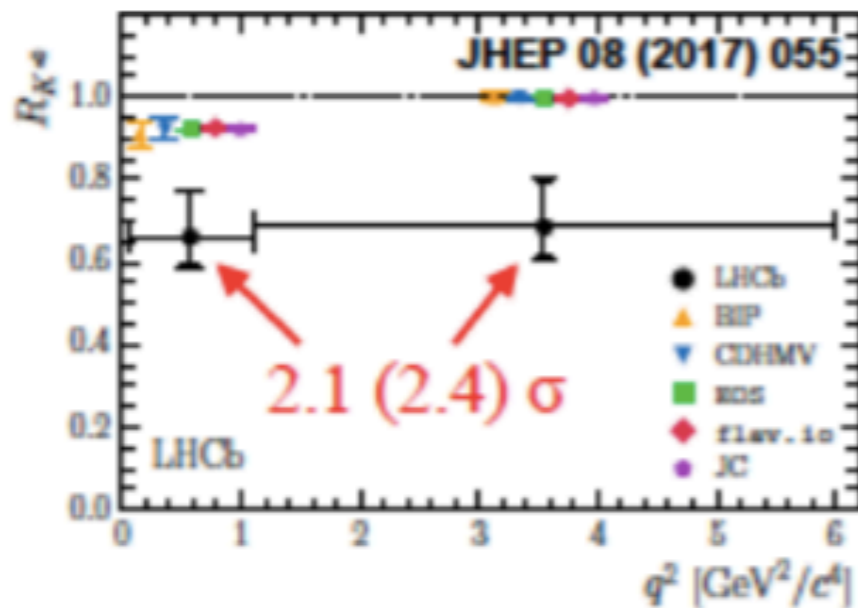
$R(\mathcal{A}_c) = 0.242 \pm 0.026 \pm 0.040 \pm 0.059(\text{ext})$

arxiv:2201:03497

Measurement is consistent with SM ($\sim 1 \sigma$ “low”) [SM= 0.324 ± 0.004].

FCNC - SM loop process: $R_{K^{(*)}}$ anomaly

$$R_{K^{(*)}} = \left. \frac{BR(B \rightarrow K^{(*)} \mu \mu)}{BR(B \rightarrow K^{(*)} e e)} \right|_{q^2 \in [q_{min}^2, q_{max}^2]}$$

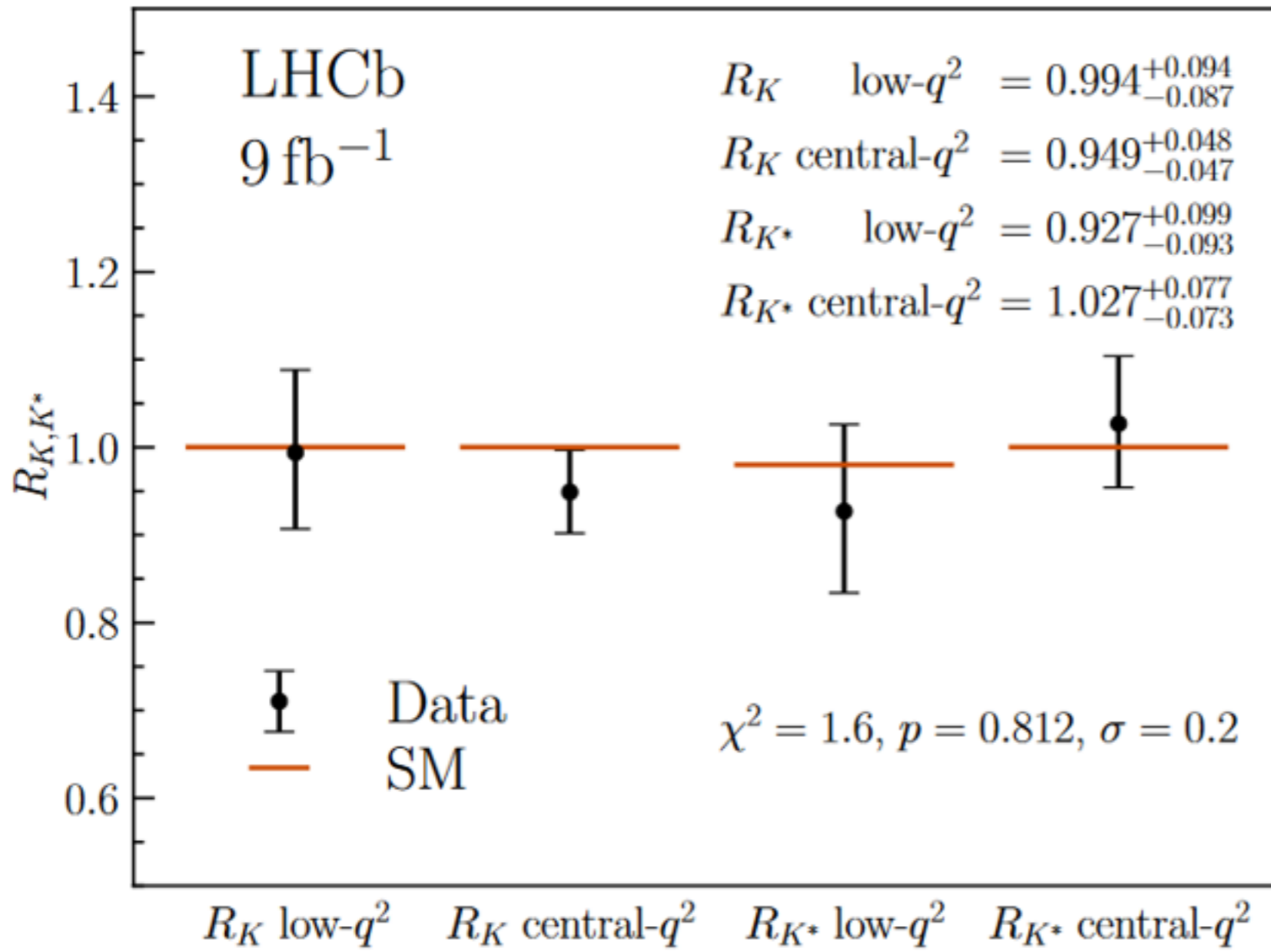


P_5' in $B \rightarrow K^* \mu^+ \mu^-$ (angular distribution functions) 3σ

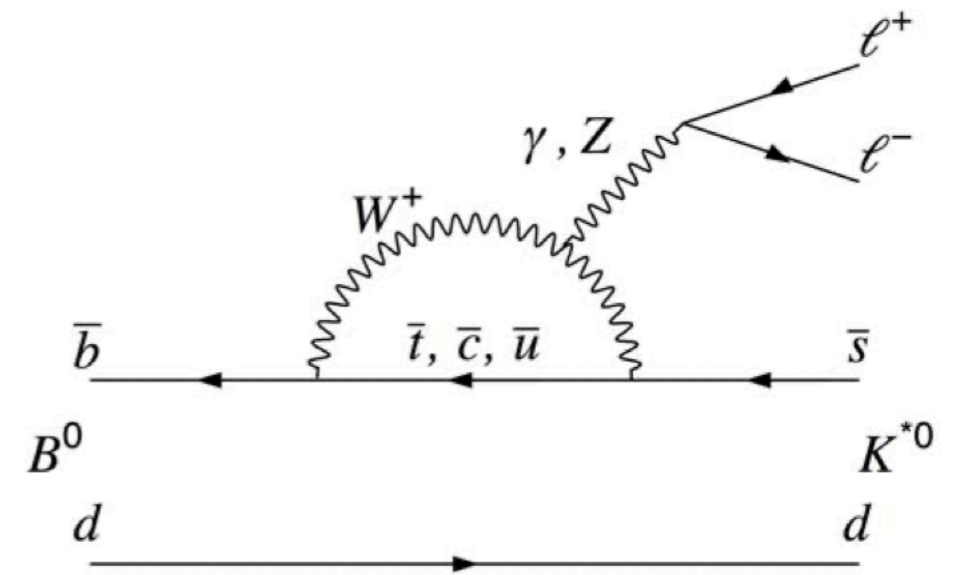
LHCb: the discrepancy present in $B_s \rightarrow \phi \mu \mu$ and $\Lambda_b \rightarrow \Lambda \mu \mu$

(see Capriotti talk at ICHEP2018)

THE LHCb RECENT RESULTS



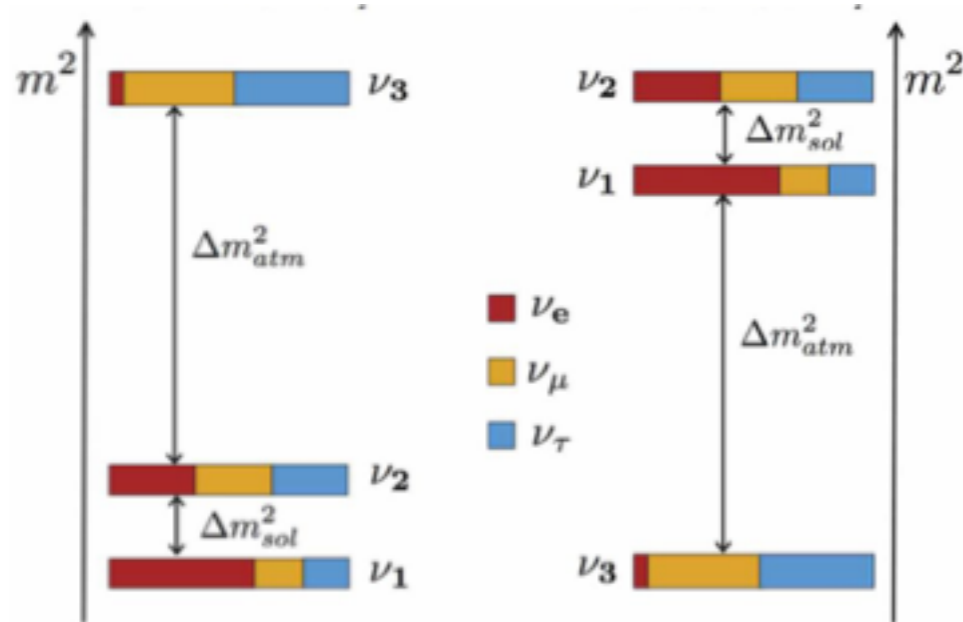
$$R_{K^{(*)}} = \frac{BR(B \rightarrow K^{(*)} \mu \mu)}{BR(B \rightarrow K^{(*)} ee)}$$



$$\mathcal{R}(K^*)_{SM} = 1.0$$

Neutrinos

Neutrino Physics



- Absolute value of neutrino masses ?
- Mass hierarchy?
- Dirac or Majorana?
- Fourth sterile neutrino?
- Neutrino dark matter?

$$0.06 \text{ eV} < \sum m_\nu < 0.12 \text{ eV}$$

↑
↑
ν-OSC
CMB

PMNS-matrix parameters are measured with high accuracy of few %

- Normal hierarchy favoured at 3.1σ
- Nonzero CP phase favoured
- Upper octant favoured

parameter	best fit $\pm 1\sigma$	3σ range
Δm_{21}^2 [10^{-5}eV^2]	$7.55^{+0.20}_{-0.16}$	7.05–8.14
$ \Delta m_{31}^2 $ [10^{-3}eV^2] (NO)	2.50 ± 0.03	2.41–2.60
$ \Delta m_{31}^2 $ [10^{-3}eV^2] (IO)	$2.42^{+0.03}_{-0.04}$	2.31–2.51
$\sin^2 \theta_{12} / 10^{-1}$	$3.20^{+0.20}_{-0.16}$	2.73–3.79
$\sin^2 \theta_{23} / 10^{-1}$ (NO)	$5.47^{+0.20}_{-0.30}$	4.45–5.99
$\sin^2 \theta_{23} / 10^{-1}$ (IO)	$5.51^{+0.18}_{-0.30}$	4.53–5.98
$\sin^2 \theta_{13} / 10^{-2}$ (NO)	$2.160^{+0.083}_{-0.069}$	1.96–2.41
$\sin^2 \theta_{13} / 10^{-2}$ (IO)	$2.220^{+0.074}_{-0.076}$	1.99–2.44
δ / π (NO)	$1.32^{+0.21}_{-0.15}$	0.87–1.94
δ / π (IO)	$1.56^{+0.13}_{-0.15}$	1.12–1.94

Is it just the SM or requires New physics?

Three Types of Seesaw Mechanisms

Require the existence of new degrees of freedom (particles) beyond those present in the SM

Type I seesaw mechanism: ν_{IR} - RH ν s' (heavy).

Type II seesaw mechanism: $\mathbf{H}(x)$ - a triplet of H^0, H^-, H^{--} Higgs fields.

Type III seesaw mechanism: $\mathbf{T}(x)$ - a triplet of fermion fields.

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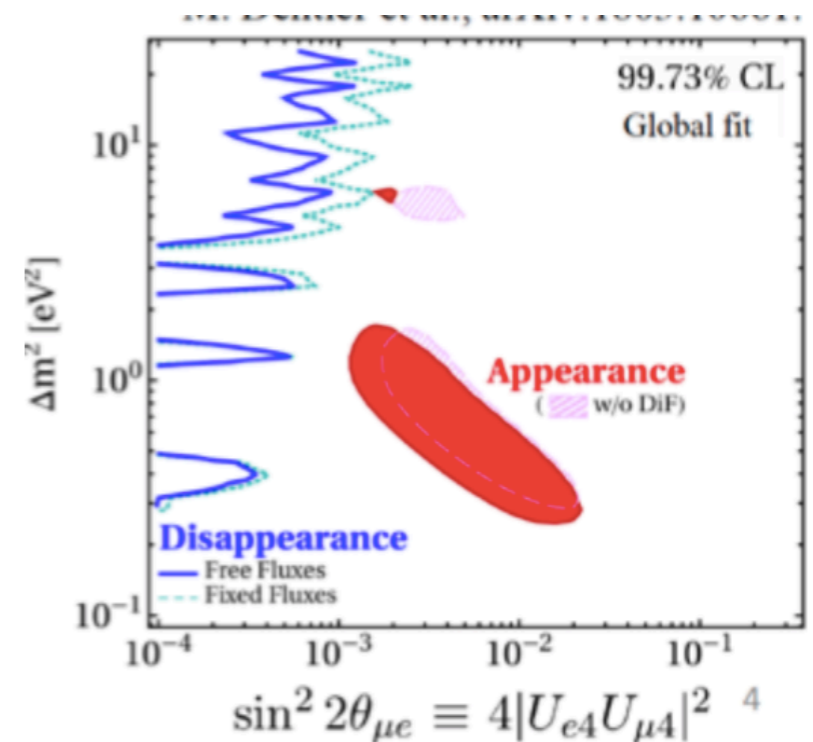
Type II seesaw mechanism: $H(x)$ - a triplet of H^0, H^-, H^{--} Higgs fields.

Type III seesaw mechanism: $T(x)$ - a triplet of fermion fields.

M. Weber ICHEP2018

• Possible Sterile Neutrino?

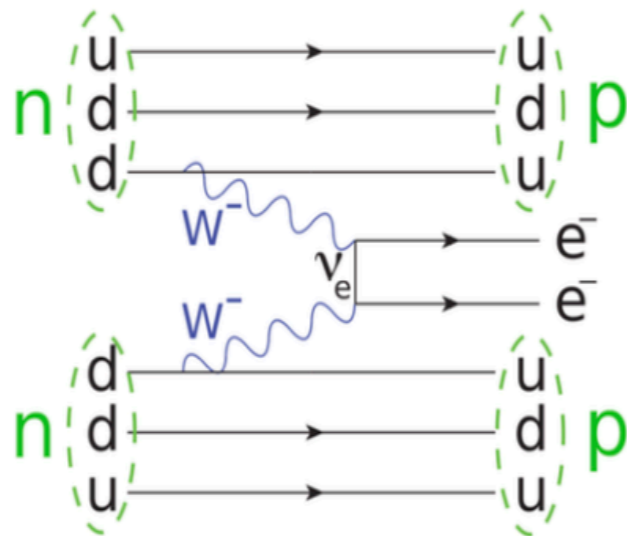
- **New MiniBooNE consistent with LSND (but low energy excess?)**
- **Reactor anomaly questioned by Daya Bay/RENO time dependence**
- **New SBL and source experiments**
- **Conflict with ν_μ disappearance**



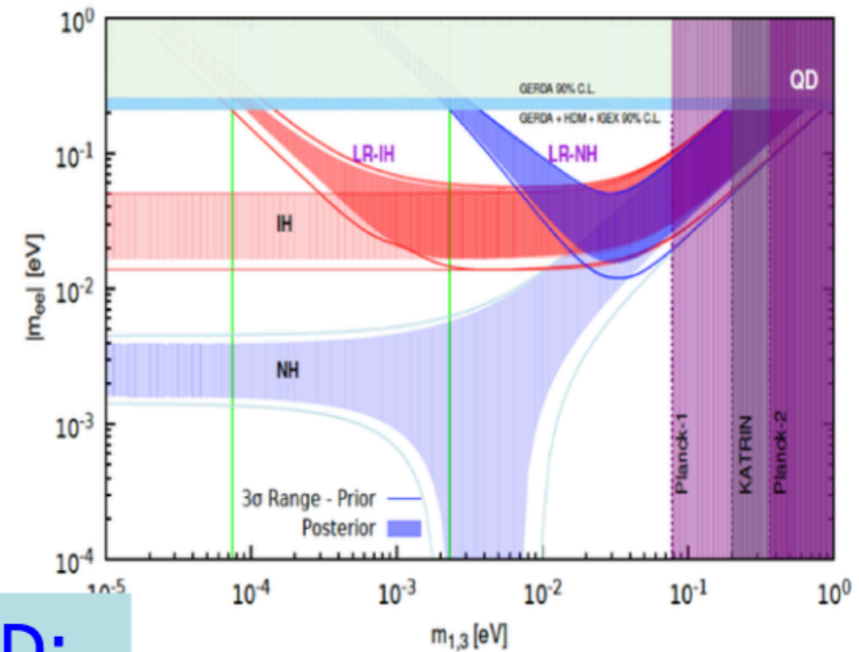
Neutrinoless Double Beta Decay

The question is still unanswered:

Are neutrinos their own antiparticles?



Ton scale 0νββ experiments will cover the inverted hierarchy by 2035



Many experiments operating, planned or in R&D:
 LEGEND, SNO+, NEXT, CUPID, THEIA...

Sterile Neutrinos

Several anomalies around in the community since some years...
Additional sterile neutrinos as a possible candidate explanation

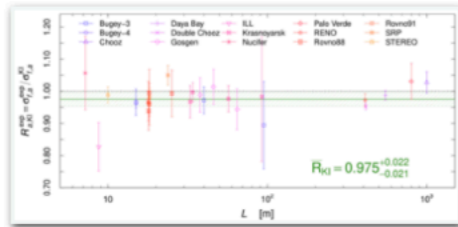
- ☑ Very generic extension of SM
 - can be leftover of extended gauge multiplet
 - ☑ Useful phenomenological tool
 - can explain ν masses (seesaw mechanism, $m \sim \text{TeV} \dots M_{\text{Pl}}$)
 - can explain cosmic baryon asymmetry (leptogenesis, $m \gg 100 \text{ GeV}$)
 - can explain dark matter ($m \sim \text{keV}$)
 - can explain oscillation anomalies ($m \sim \text{eV}$)
- Promote mixing matrix to 4×4 , oscillation formula unchanged:



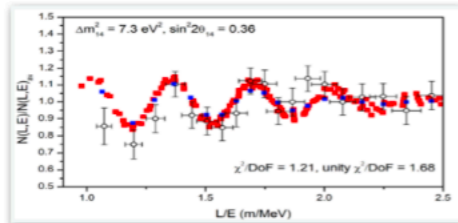
$$P_{\alpha \rightarrow \beta} = \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* \exp \left[-i(E_j - E_k)T \right]$$

Neutrino Anomalies

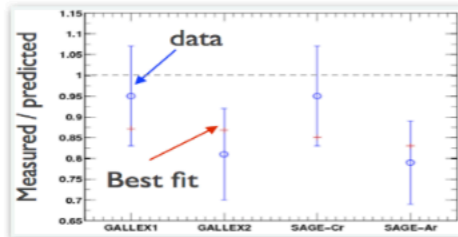
Albert De Roeck



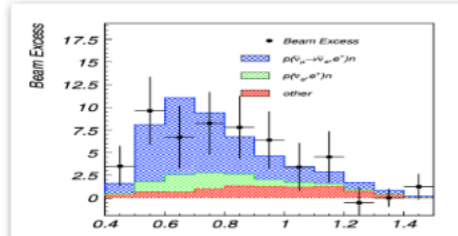
reactor flux anomaly
resolved with new input data
to flux calculation



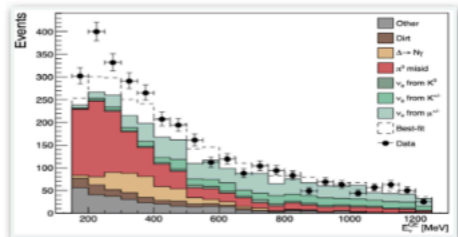
reactor spectra
is there really an anomaly? -> DANSS



gallium anomaly
unresolved, recently reinforced BEST



LSND
unresolved



MiniBooNE
unresolved μ BooNe excluded some explanations
resolvable by next-gen. SBL experiments



More details in the backup

- Jury still out on many of these anomalies. No clear picture emerging yet.
- Simple sterile neutrino would not fit all the data. Tensions on all sides...
- Future: Reactor experiments continuing or new ones (eg JSNS²) or new experiments at the FNAL short neutrino baseline... (ICARUS, SBND)

BEYOND THE STANDARD MODEL: THE MASS SPECTRUM AND MIXINGS

- Mass spectrum?

$$m_{quark} = y_{quark} \cdot v$$

$$m_{lepton} = y_{lepton} \cdot v$$

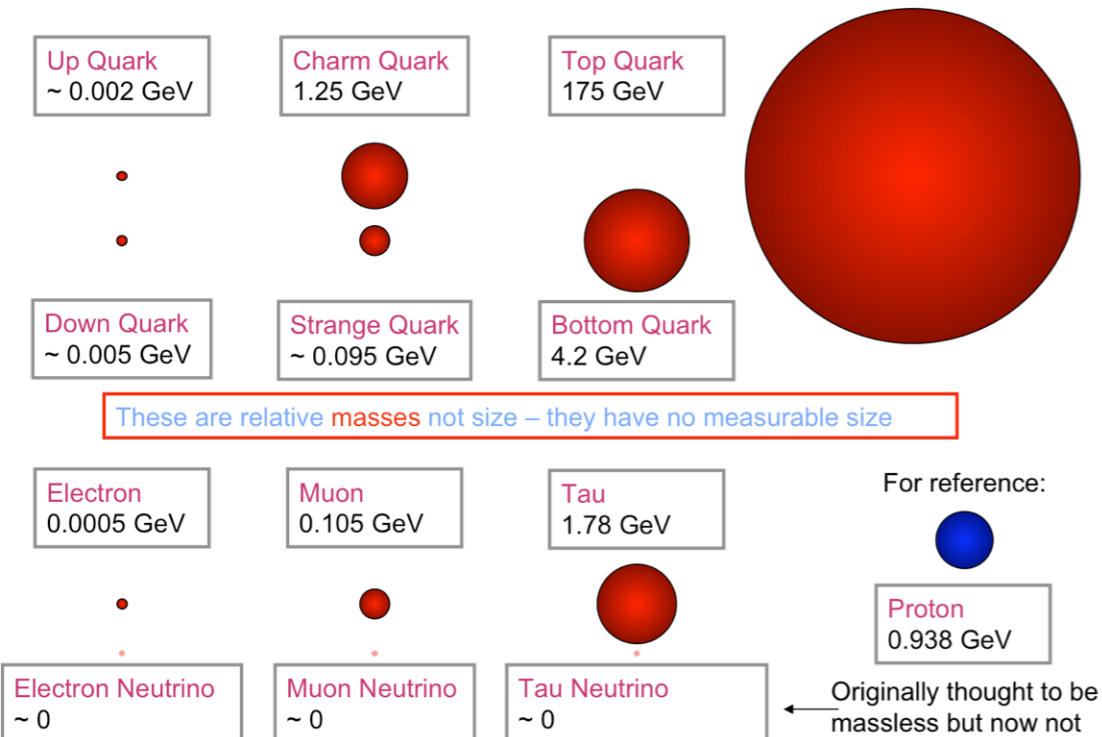
$$m_W = g/\sqrt{2} \cdot v$$

$$m_Z = \sqrt{g^2 + g'^2}/\sqrt{2} \cdot v$$

$$m_H = \sqrt{\lambda} \cdot v$$

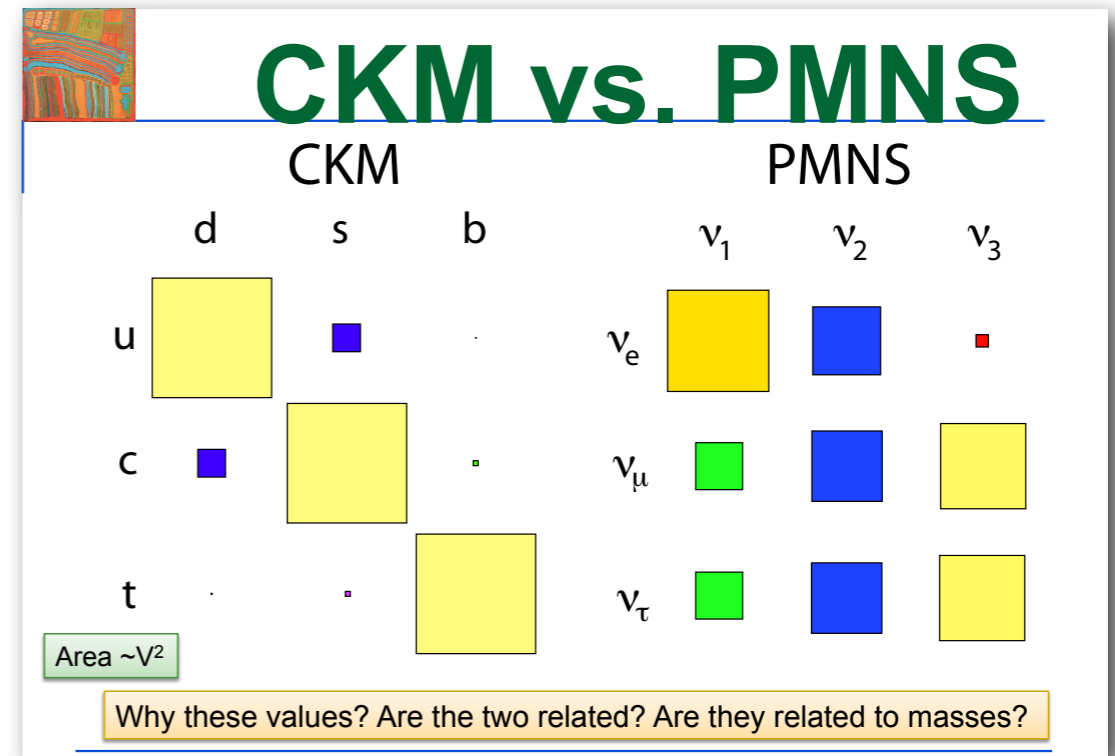
SM $m_\gamma = 0$

$$m_{gluon} = 0$$



- Mixing Matrices?

- Quark-Lepton Symmetry
- Strong difference in parameters



- What are the CKM and PMNS phases?
- Where lies the source of CP violation: in quark or lepton sector?

$$J_{CP} = \frac{1}{8} \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta$$

Neutrino Mixing: New Symmetry?

- $\theta_{12} = \theta_{\odot} \simeq \frac{\pi}{5.4}$, $\theta_{23} = \theta_{\text{atm}} \simeq \frac{\pi}{4}(?)$, $\theta_{13} \simeq \frac{\pi}{20}$

$$U_{\text{TBM}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \end{pmatrix}; \quad U_{\text{BM}} = \begin{pmatrix} \frac{1}{\sqrt{2}} & \pm \frac{1}{\sqrt{2}} & 0 \\ -\frac{1}{2} & \pm \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{2} & \mp \frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix}. \quad U_{\text{HGM}} = \begin{pmatrix} \frac{\sqrt{3}}{2} & \frac{1}{2} & 0 \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

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Follows the attempts in quark sector with 30 years delay: so far unsuccessful

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Follows the attempts in quark sector with 30 years delay: so far unsuccessful

$$U_{\text{TBM}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \end{pmatrix} \quad U_{\text{HGM}} = \begin{pmatrix} \pm \frac{1}{\sqrt{2}} & 0 \\ -\frac{1}{2} & \pm \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{2} & \mp \frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad U_{\text{HGM}} = \begin{pmatrix} \frac{\sqrt{3}}{2} & \frac{1}{2} & 0 \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ -\frac{1}{2\sqrt{2}} & \frac{\sqrt{3}}{2\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$



Neutrino Mixing: New Symmetry?

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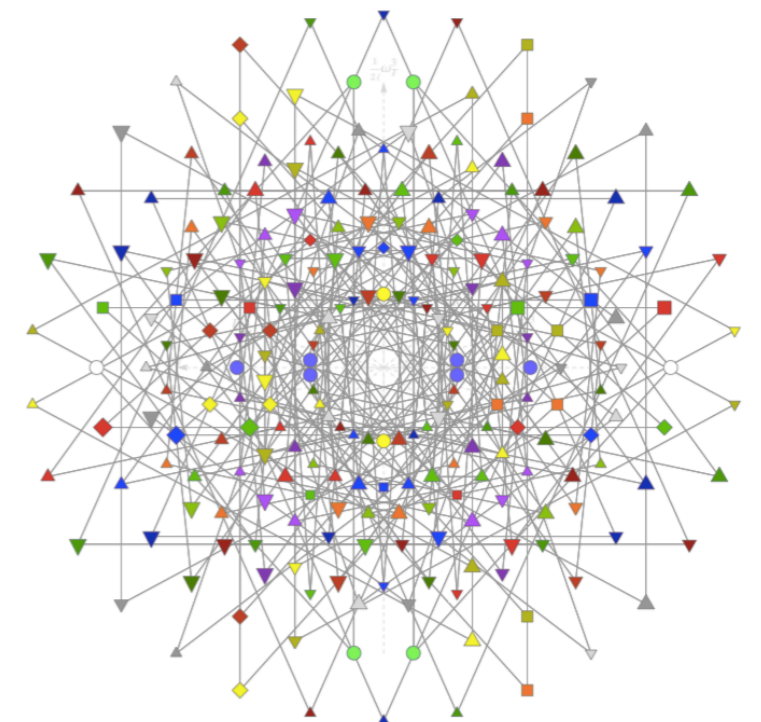


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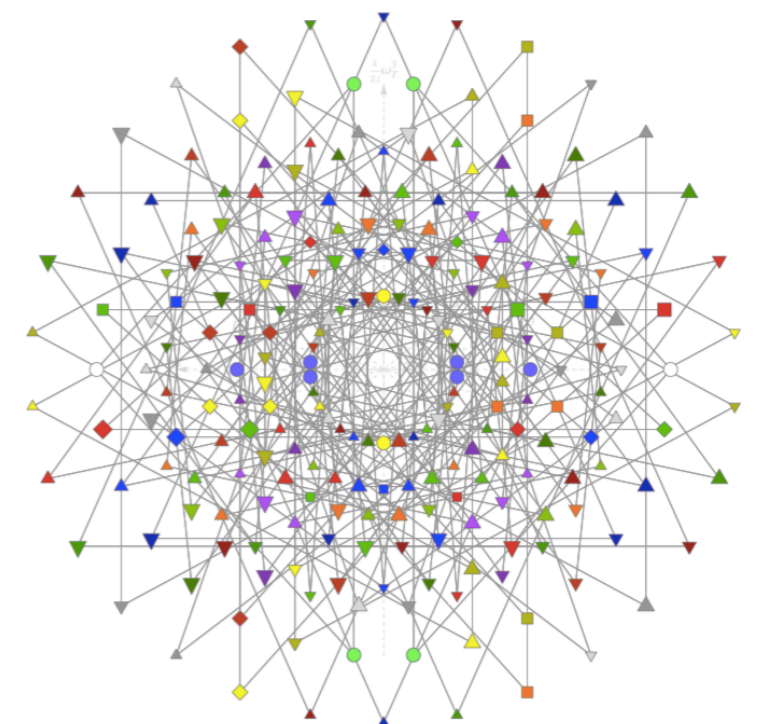
E8 roots

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Symmetry might be tricky

E8 roots

Dark matter

Major problem: 85% of matter is dark and remains invisible!

Is this compatible with the SM?

Does it requires modification of the SM or addition of gravity?

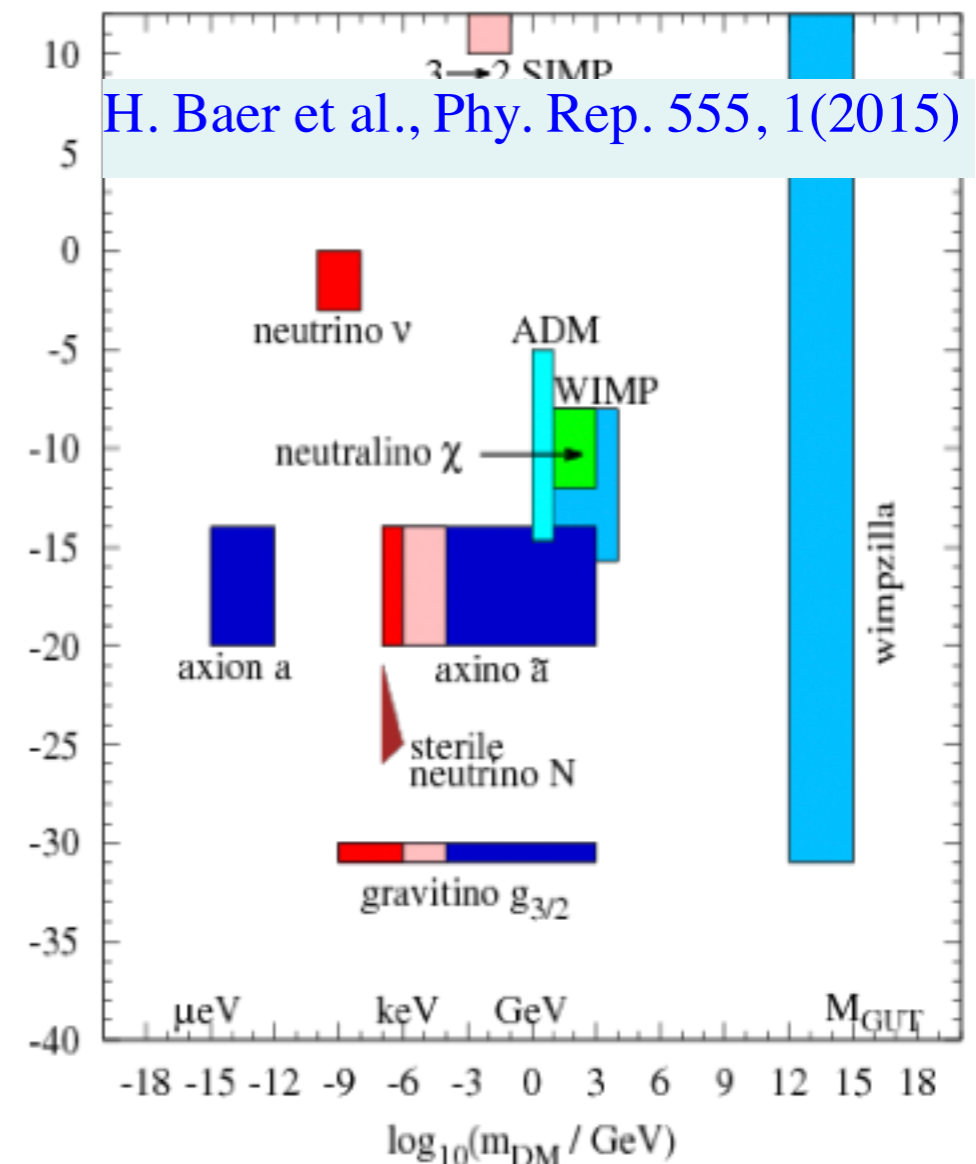
- Many candidates in many orders of magnitude of mass:

- **MOND** (Problems: large scales, Bullet cluster)
- **Primordial black holes** (LIGO, but constraints)
- **Fuzzy** (very light bosons)
- **Warm** (KeV sterile)
- **WIMP**
- **Axions/ALPs**
- **Dark sector**
- **Gravitinos**
- **Moduli**
- **Wimpzillas**



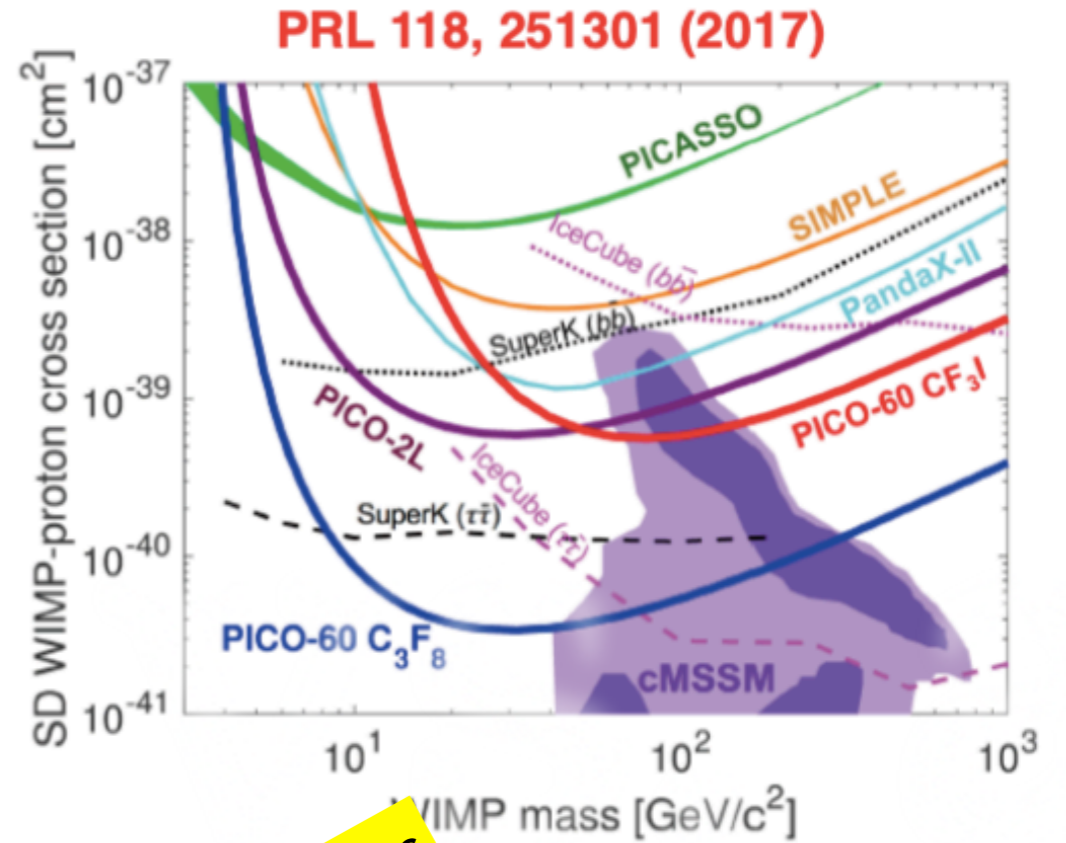
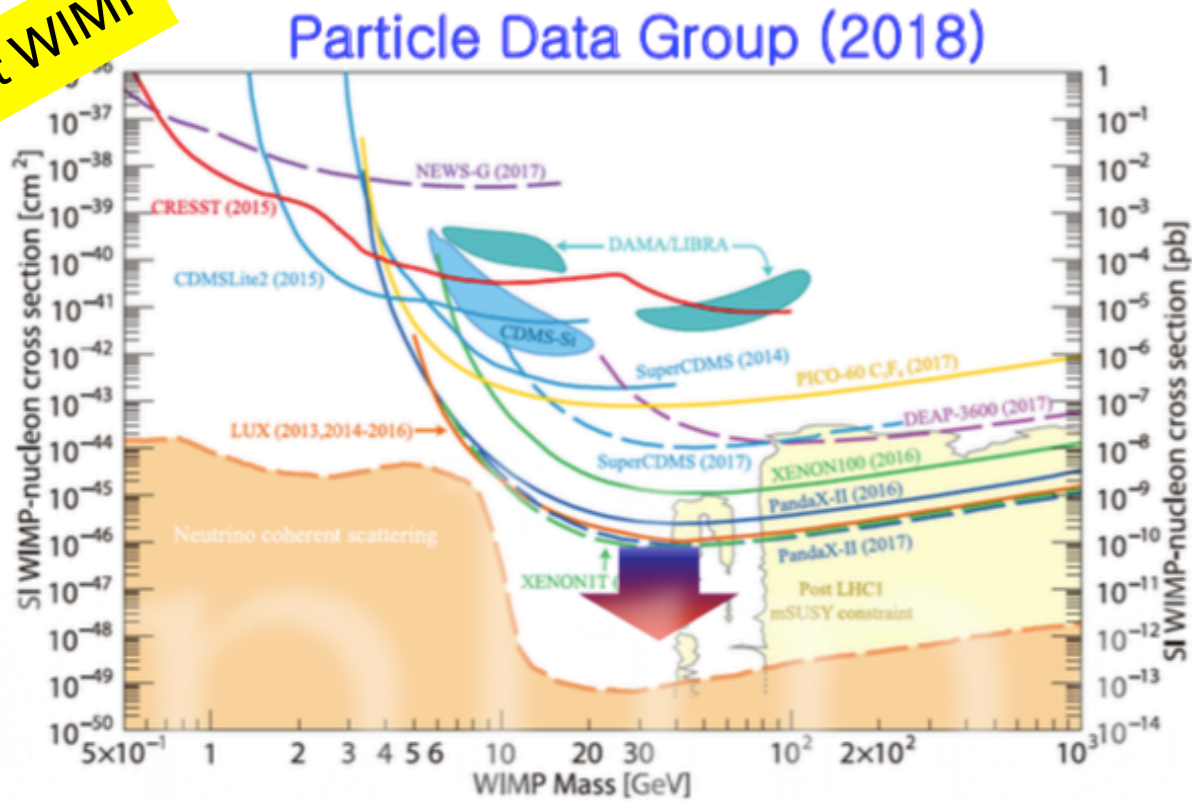
M. Drees

- Direct, indirect, collider

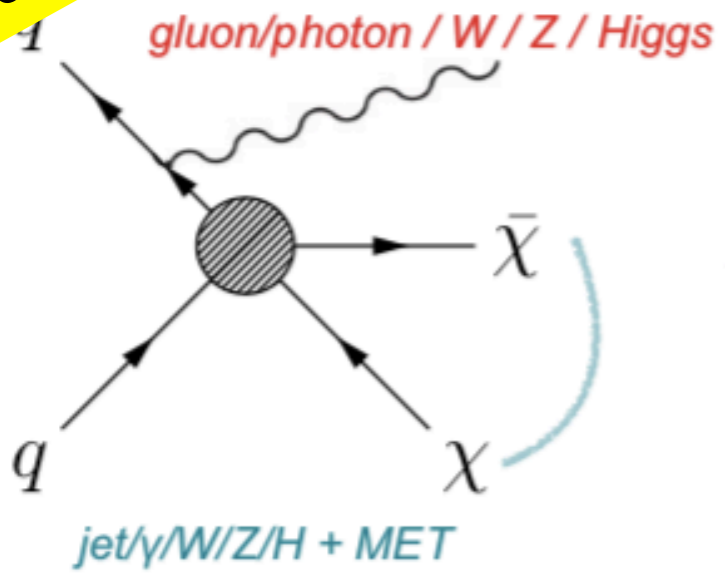


BEYOND THE STANDARD MODEL: DARK MATTER SEARCHES

Direct WIMP



Colliders WIMP

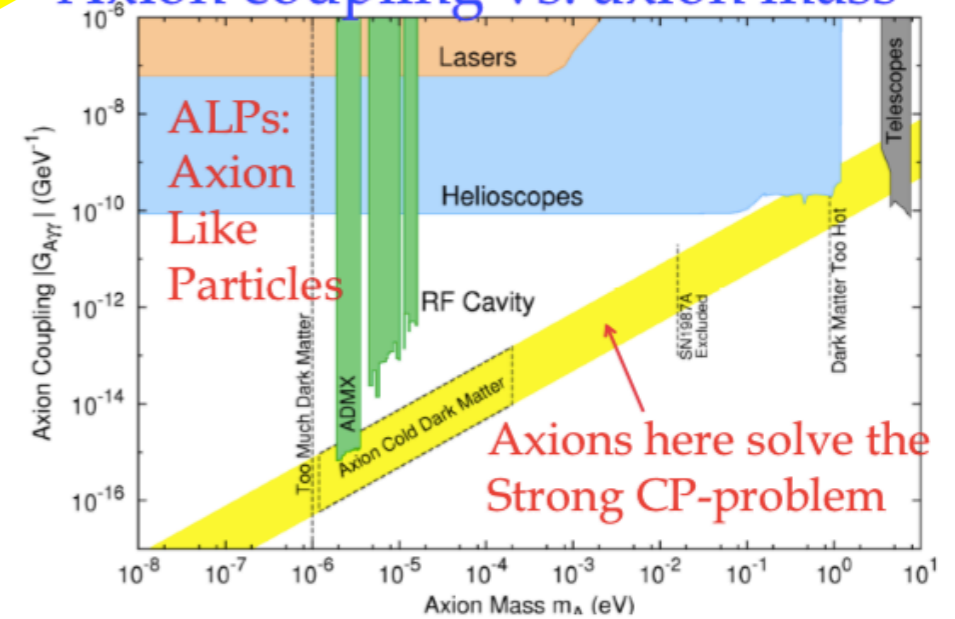


- **mono-jet**
 - most general signature, constraints on many models
- **mono-photon**
 - more challenging for background estimation
 - less powerful: EW vs. strong interaction
- **mono-W/Z leptonic**
 - clean signature and simple trigger
 - penalized by W/Z branching fraction
- **mono-W/Z hadronic**
 - larger statistics with larger background
- **tt+MET/bb+MET and mono-top**
 - more complicated experimentally
 - powerful in some scenarios
- **mono-Higgs**
 - powerful in some scenarios

D. del Re

Axion-likes

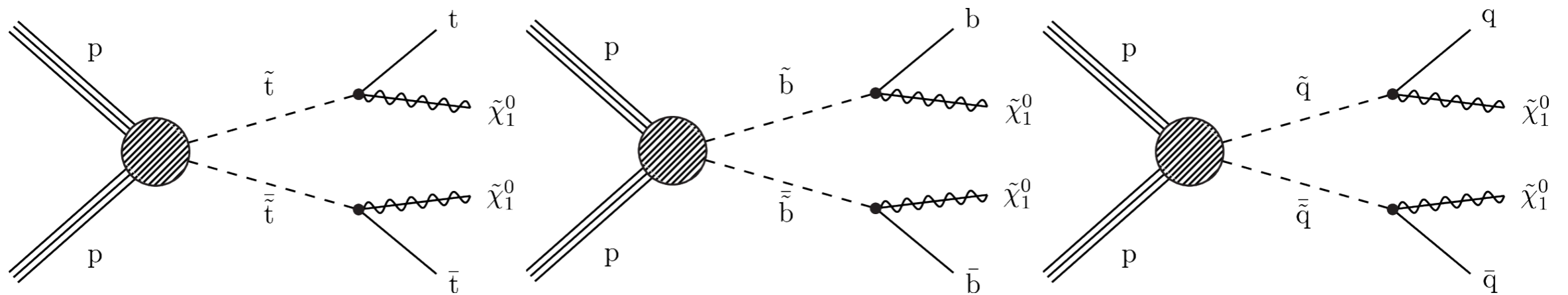
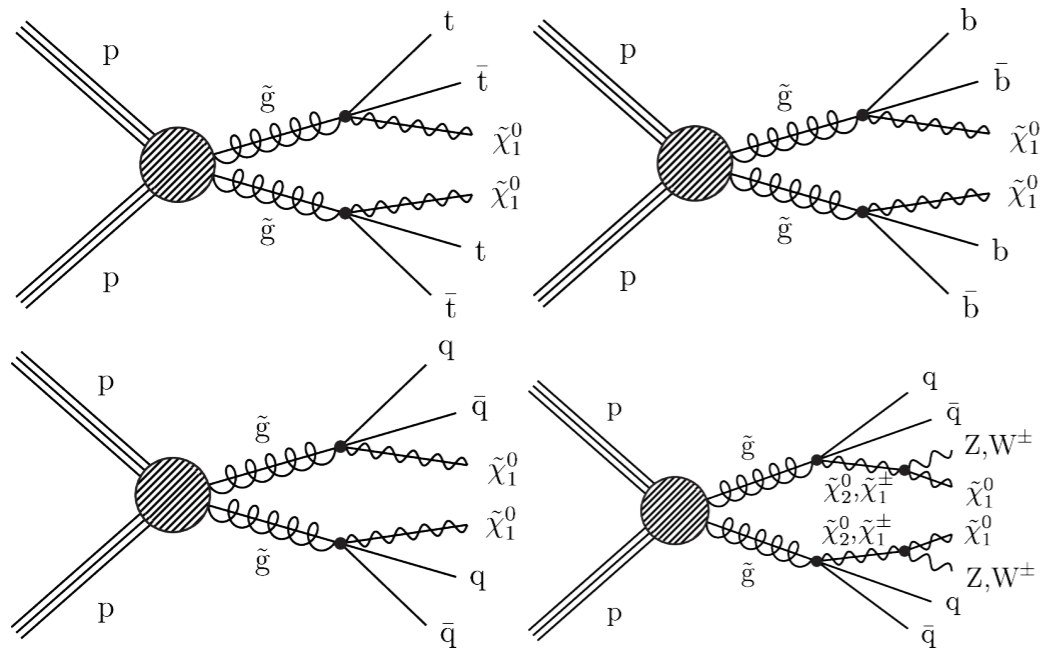
Axion coupling vs. axion mass



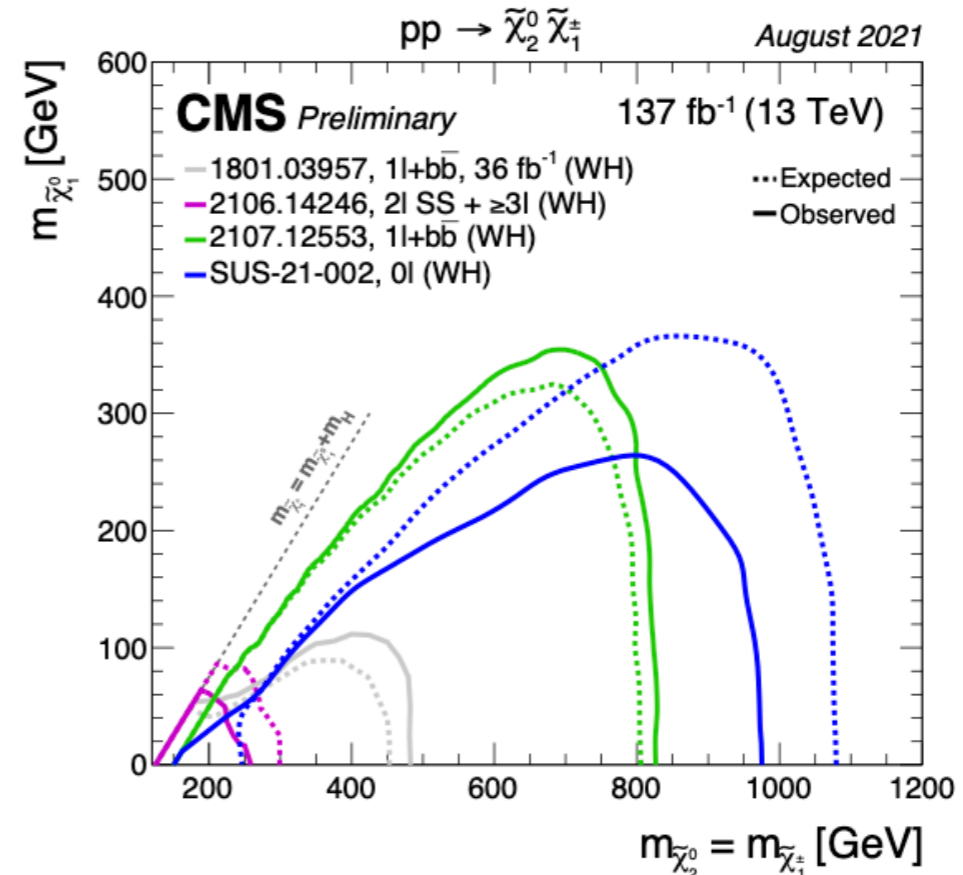
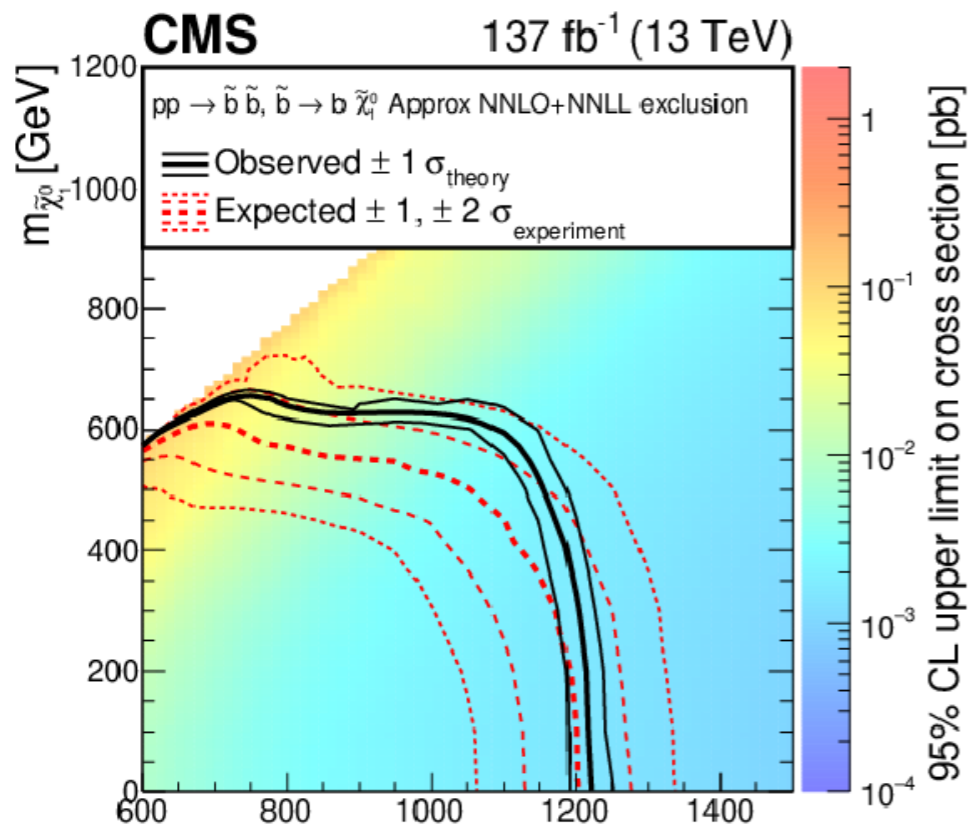
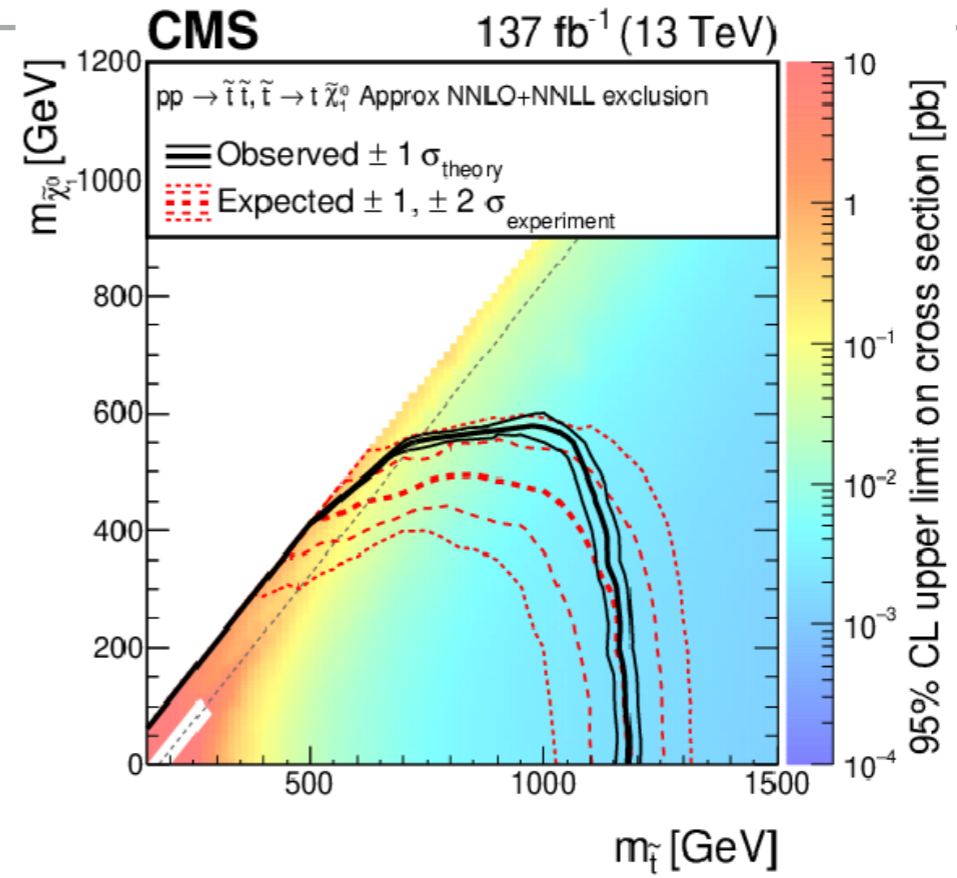
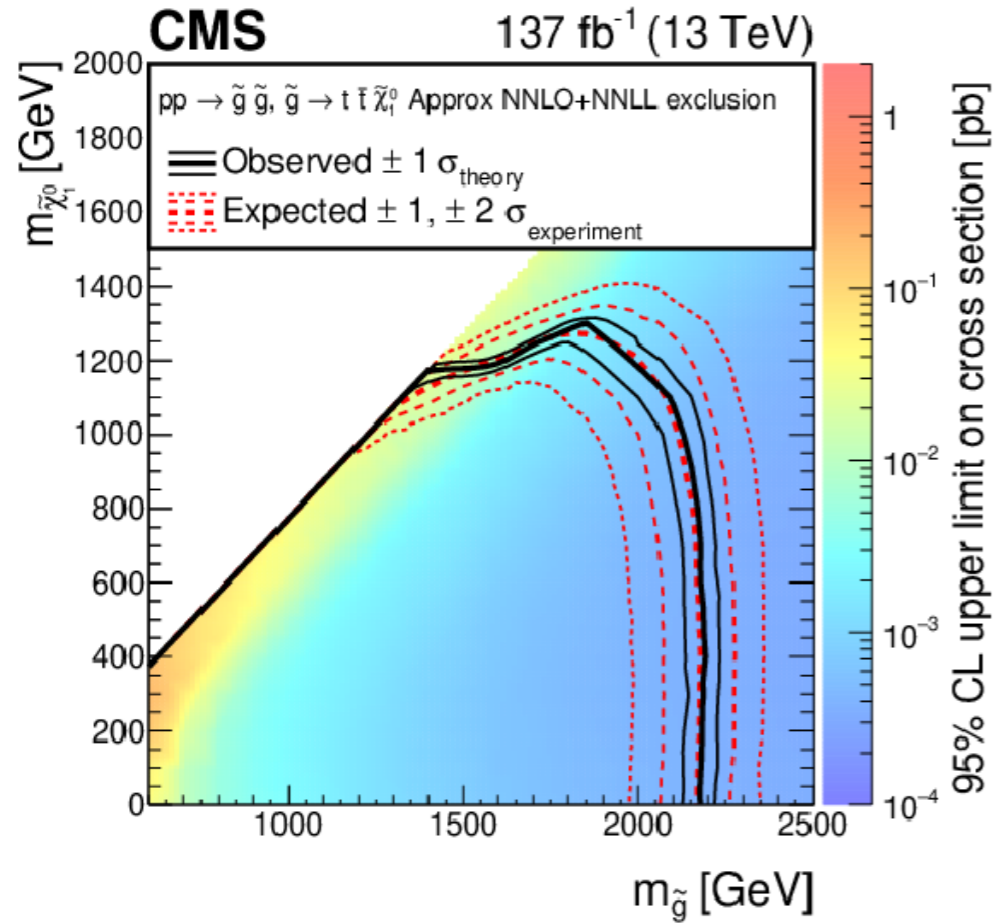
Y. Semertzidis

SUSY Searches

SUSY PRODUCTION

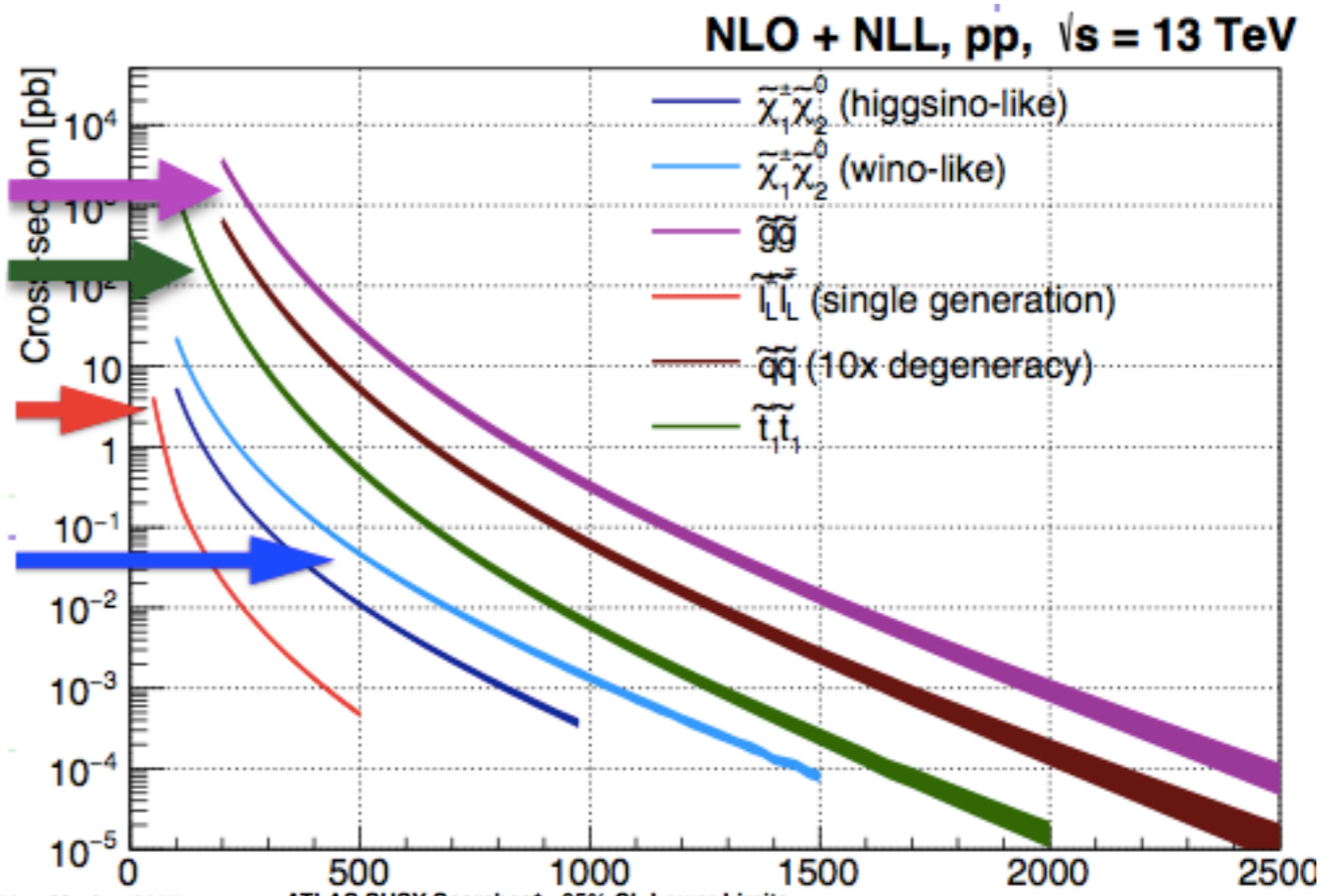


SUSY LIMITS



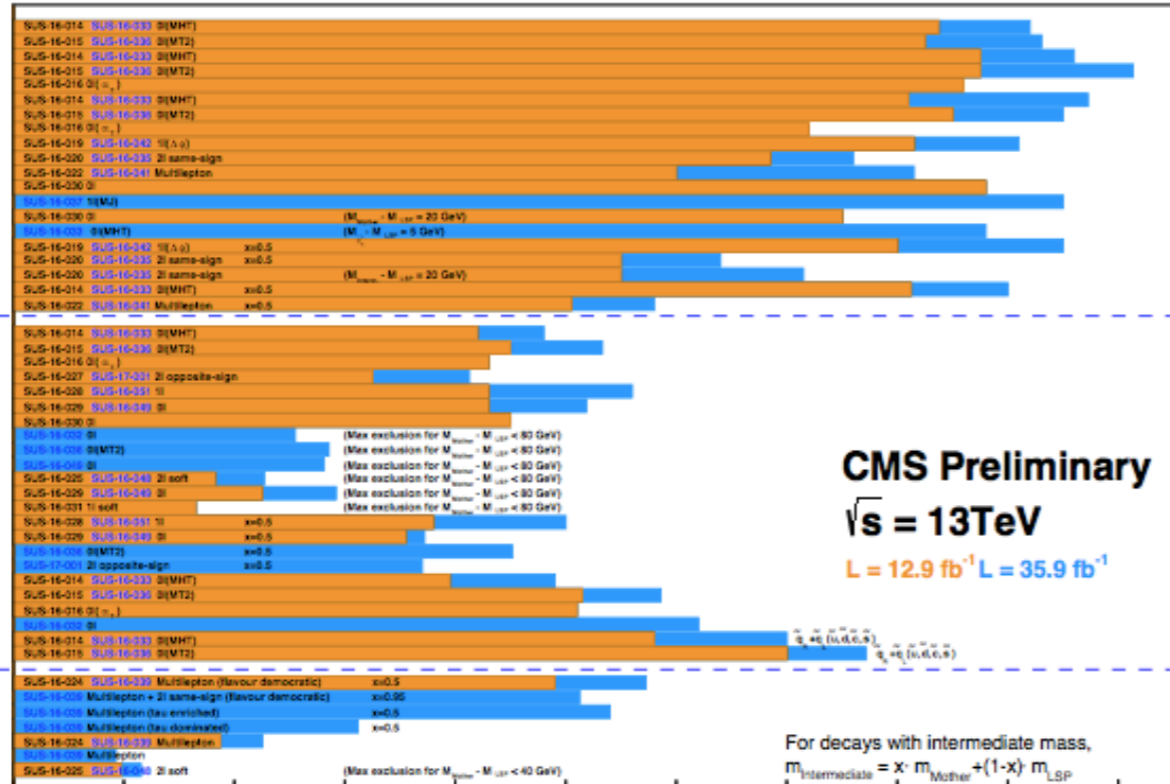
Strong production
(gluinos, squarks)

EWK production
(charginos, neutralinos, sleptons)



Selected CMS SUSY Results* - SMS Interpretation

ICHEP '16 - Moriond '17



*Observed limits at 95% C.L. - theory uncertainties not included
Only a selection of available mass limits. Probe *up to* the quoted mass limit for $m_{\text{LSP}} = 0$ GeV unless stated otherwise

ATLAS SUSY Searches* - 95% CL Lower Limits
May 2017

ATLAS Preliminary
 $\sqrt{s} = 7, 8, 13$ TeV

Model	$\epsilon, \mu, \tau, \gamma$	Jets	$E_{\text{miss}}^{\text{min}}$	$f_{\text{cut}}^{\text{min}}$	Mass limit	$\sqrt{s} = 7, 8$ TeV	$\sqrt{s} = 13$ TeV	Reference
Inclusive Searches	MDG/GRACMSM	0-3 $\epsilon, \mu, 1-2 + 2$ jets	Yes	20.3	5.85 TeV	$m_{\tilde{g}} = 1000$ GeV	$m_{\tilde{g}} = 1000$ GeV	1507.0525
1/2 gen. squarks (no top)	$\tilde{u}, \tilde{d} \rightarrow q\bar{q}$	0	2-4 jets	Yes	36.1	506 GeV	1.97 TeV	1504.0773
	$\tilde{u}, \tilde{d} \rightarrow q\bar{q}$ (compressed)	mono-jet	1-3 jets	Yes	32.2		2.02 TeV	1507.0593
	$\tilde{u}, \tilde{d} \rightarrow q\bar{q} + \gamma$	0	2-4 jets	Yes	36.1		2.01 TeV	1507.0593
	$\tilde{u}, \tilde{d} \rightarrow q\bar{q} + \gamma + \gamma$	3 ϵ, μ	4 jets	Yes	36.1		1.825 TeV	1507.0593
	$\tilde{u}, \tilde{d} \rightarrow q\bar{q} + \gamma + \gamma$	2 ϵ, μ	4 jets	Yes	36.1		1.8 TeV	1507.0593
	$\tilde{u}, \tilde{d} \rightarrow q\bar{q} + \gamma + \gamma$	1-2 $\epsilon + 0-1 \gamma$	0-2 jets	Yes	32.2		2.0 TeV	1507.0593
	GGM (bino NLSP)	2 γ	2 jets	Yes	32.2		1.68 TeV	1506.09150
	GGM (higgsino bino NLSP)	7	2 jets	Yes	20.3		1.37 TeV	1507.0593
	GGM (higgsino NLSP)	2 ϵ, μ (2)	2 jets	Yes	20.3		1.8 TeV	1503.0236
	Gravitino LSP	0	mono-jet	Yes	20.3	300 GeV	885 GeV	1502.01518
1/2 gen. squarks (with top)	$\tilde{u}, \tilde{d} \rightarrow q\bar{q}$	0	3-5 jets	Yes	36.1		1.92 TeV	1507.0593
	$\tilde{u}, \tilde{d} \rightarrow q\bar{q} + \gamma$	0.1 ϵ, μ	3-5 jets	Yes	36.1		1.97 TeV	1507.0593
	$\tilde{u}, \tilde{d} \rightarrow q\bar{q} + \gamma$	0.1 ϵ, μ	3-5 jets	Yes	36.1		1.37 TeV	1407.0803
1/2 gen. squarks (with top)	$\tilde{u}, \tilde{d}, \tilde{t}, \tilde{b} \rightarrow q\bar{q}$	0	2-5 jets	Yes	36.1	950 GeV		1507.0593
	$\tilde{u}, \tilde{d}, \tilde{t}, \tilde{b} \rightarrow q\bar{q}$	2 ϵ, μ (50)	1-5 jets	Yes	36.1	113-119 GeV	275-700 GeV	1507.0593
	$\tilde{u}, \tilde{d}, \tilde{t}, \tilde{b} \rightarrow q\bar{q} + \gamma$	0.2 ϵ, μ	1-2-5 jets	Yes	20.3/26.1	90-790 GeV	255-690 GeV	1508.08616, 1508.08616, 1508.08616
	$\tilde{u}, \tilde{d}, \tilde{t}, \tilde{b} \rightarrow q\bar{q} + \gamma$	0	mono-jet	Yes	32.2	90-323 GeV	150-600 GeV	1504.0773
EW direct	$\tilde{g}, \tilde{g} \rightarrow q\bar{q}$	2 ϵ, μ	0	Yes	36.1	89-440 GeV		1507.0593
	$\tilde{g}, \tilde{g} \rightarrow q\bar{q} + \gamma$	2 ϵ, μ	0	Yes	36.1	710 GeV		1507.0593
	$\tilde{g}, \tilde{g} \rightarrow q\bar{q} + \gamma + \gamma$	2 ϵ, μ	0	Yes	36.1	780 GeV		1507.0593
	$\tilde{g}, \tilde{g} \rightarrow q\bar{q} + \gamma + \gamma$	2 ϵ, μ	0	Yes	36.1	545 GeV	1.18 TeV	1507.0593
	$\tilde{g}, \tilde{g} \rightarrow q\bar{q} + \gamma + \gamma$	0.3 ϵ, μ	0-2 jets	Yes	36.1	270 GeV	580 GeV	1507.0593
Long-lived particles	Direct $\tilde{g}, \tilde{g} \rightarrow q\bar{q}$, long-lived \tilde{g}	disapp. tk	1 jet	Yes	36.1	430 GeV		1507.0593
	Direct $\tilde{g}, \tilde{g} \rightarrow q\bar{q}$, long-lived \tilde{g}	dChalk	-	Yes	18.4	495 GeV	850 GeV	1505.0532
	Stable, stopped \tilde{g} R-hadron	0	1-5 jets	Yes	27.9			1310.0594
	Stable \tilde{g} R-hadron	ok	-	-	3.2			1504.0420
RPV	Metastable \tilde{g} R-hadron	dChalk	-	-	3.2			1504.0420
	GMSB, stable $\tilde{g}, \tilde{g} \rightarrow q\bar{q} + \gamma + \gamma$	1-2 ϵ, μ	-	Yes	20.3	337 GeV		1411.6795
	GMSB, $\tilde{g}, \tilde{g} \rightarrow q\bar{q}$, long-lived \tilde{g}	2 γ	-	Yes	20.3	440 GeV		1409.5542

Baryon Asymmetry of the Universe

Baryon Asymmetry of the Universe



SM expectation:

$$\frac{n_B - n_{\bar{B}}}{n_\gamma} \sim 10^{-18}$$

vs.

Observed*:

$$\frac{n_B - n_{\bar{B}}}{n_\gamma} \sim 10^{-10}$$

Sakharov criteria

1. Baryon number violation
2. C and CP violation
3. Thermal non-equilibrium



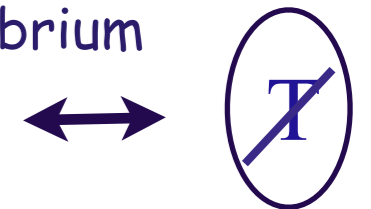
*WMAP


THE STANDARD MODEL: CONCEPTUAL PROBLEMS

- Baryon asymmetry of the Universe

1. Violation of a thermal equilibrium

A possible scenario in the early Universe when particles drop from thermal equilibrium
violations T invariance

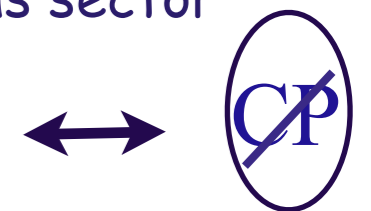


2. Violation of baryon number \longleftrightarrow  $B = \frac{N_q - N_{\bar{q}}}{3}$

Baryon number is conserved in the SM with exponential accuracy

Violation of baryon number occurs in Grand Unified Theories and in Lepton=fourth color models (Pati-Salam model) } New particles = Leptoquarks, Extended Higgs sector

3. Violation of CP invariance (requires larger CP than in the SM)



In the SM achieved via phase factors in the CKM and PMNS mixing matrices

The presence of new phase factors in extended models (2HDM, SUSY, etc)

CPT is exact symmetry of Nature

BEYOND THE STANDARD MODEL: CONCLUSIONS

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- Lack of understanding of flavor structure of the SM calls for explanation at higher level
- New era in gravity due to discovery of gravitational waves and black holes might change the landscape

Ideas (conventional and not)

- **Symmetries**
 - Supersymmetry, family, ...
- **Compositeness**
 - Higgs, fermions, ...
- **Extra dimensions**
 - large, warped, ...
- **Dark or hidden sectors**
 - Dark, SUSY-breaking, random, ...
- **Unification**
 - GUT, string, ...
- **New dynamical ideas**
 - Relaxion, unnaturalness, clockwork, string instantons, ...
- **Random or environmental**
 - multiverse
- **String remnants**
 - (need not solve SM problem)
 - Z' , vector fermions, extended Higgs, dark, moduli, axions, ...

BEYOND THE STANDARD MODEL: CONCLUSIONS



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BEYOND THE STANDARD MODEL: CONCLUSIONS



How Will We Make Progress?

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-  **The energy frontier**
-  **The intensity frontier**
-  **The precision frontier**
-  **Underground and neutrino**
-  **Cosmology and astrophysics**

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